## Atmosphere, Ocean and Climate Dynamics Answers to Chapter 5

1. Fig. 5.5 shows the net incoming solar and outgoing longwave irradiance at the top of the atmosphere. Note that there is a net gain of radiation in low latitudes and a net loss in high latitudes. By inspection of the figure, estimate the magnitude of the poleward energy flux that must be carried by the atmosphere-ocean system across the $30^{\circ}$ latitude circle, to achieve a steady state.
The energy absorbed between the equator and $30^{\circ}$ latitude is, roughly, $\frac{1}{2} \times 50 \mathrm{~W} \mathrm{~m}^{-2} \times \frac{1}{4} \times 5 \times 10^{14} \mathrm{~m}^{2}=3.125 \times 10^{15} \mathrm{~W}$, roughly comparable to Fig.5.6. [Note $30^{\circ}$ latitude divides the hemisphere in to equal areas.]
2. Suppose that the Earth's rotation axis were normal to the Earth-Sun line. The solar flux, measured per unit area in a plane normal to the Earth-Sun line, is $S_{0}$. By considering the solar flux incident on a latitude belt bounded by latitudes $(\varphi, \varphi+d \varphi)$, show that $F$, the $24 h r$ averaged solar flux per unit area of the Earth's surface, varies with latitude as

$$
F=\frac{S_{0}}{\pi} \cos \varphi
$$

The area of parallel solar beam intercepted by latitudes $\varphi$ and $\varphi+d \varphi$ is $2 a \cos \varphi \times d y$ where $d y=a d \varphi \cos \varphi$. The surface area of the planet between latitudes $\varphi$ and $\varphi+d \varphi$ is $2 \pi a \cos \varphi \times a d \varphi$. Thus:

$$
F=\frac{2 a \cos \varphi \times d y \times S_{0}}{2 \pi a \cos \varphi \times a d \varphi}=\frac{2 a \cos \varphi \times a d \varphi \cos \varphi \times S_{0}}{2 \pi a^{2} \cos \varphi \times d \varphi}=\frac{S_{0}}{\pi} \cos \varphi
$$

(a) Using this result, suppose that the atmosphere is completely transparent to solar radiation, but opaque to infrared such that, separately at each latitude, the radiation budget can be represented as for a "single slab" atmosphere - the "single slab" model is discussed in Section 2.3.1. Determine how surface temperature varies with latitude.
Balance at top of the atmosphere in each latitude belt:

$$
\left(1-\alpha_{p}\right) \frac{S_{0}}{\pi} \cos \varphi=A \uparrow=\sigma T_{a}^{4}
$$

Balance at ground in each latitude belt:

$$
\left(1-\alpha_{p}\right) \frac{S_{0}}{\pi} \cos \varphi+A \downarrow=\sigma T_{s}^{4}
$$

But $A \uparrow=A \downarrow=\sigma T_{a}^{4}$ hence:

$$
2\left(1-\alpha_{p}\right) \frac{S_{0}}{\pi} \cos \varphi=\sigma T_{s}^{4}
$$

and so

$$
T_{s}=\left[\frac{8 \cos \varphi}{\pi}\right]^{\frac{1}{4}}\left[\frac{\left(1-\alpha_{p}\right) S_{0}}{4 \sigma}\right]^{\frac{1}{4}}=\left[\frac{8 \cos \varphi}{\pi}\right]^{\frac{1}{4}} T_{e}
$$

where $T_{e}$ is given by Eq.(2.4).
The solution is plotted as a function of latitude below.

(b) Calculate the surface temperature at the equator, $30^{\circ}$, and $60^{\circ}$ latitude, using data for Earth albedo and $S_{0}$.

$$
\begin{aligned}
& T_{s}\left(0^{\circ}\right)=\left[\frac{8 \cos 0^{\circ}}{\pi}\right]^{\frac{1}{4}} 255=322.13 \\
& T_{s}\left(30^{\circ}\right)=\left[\frac{8 \cos 30^{\circ}}{\pi}\right]^{\frac{1}{4}} 255=310.75 \\
& T_{s}\left(60^{\circ}\right)=\left[\frac{8 \cos 60^{\circ}}{\pi}\right]^{\frac{1}{4}} 255=270.87
\end{aligned}
$$

3. Use the hydrostatic relation and the equation of state of an ideal gas to show that the $1000-500 \mathrm{mbar}$ "thickness", $\Delta z=z(500 \mathrm{mbar})-$
$z(1000 \mathrm{mbar})$ is related to the mean temperature $\langle T\rangle$ of the $1000-$ 500 mbar layer by

$$
\Delta z=\frac{R\langle T\rangle}{g} \ln 2
$$

where

$$
\langle T\rangle=\frac{\int T d \ln p}{\int d \ln p}
$$

where the integrals are from 500 mbar to 1000 mbar. (Note that $1000 \mathrm{mbar} \equiv$ $\left.1000 h \mathrm{~Pa} \equiv 10^{5} \mathrm{~Pa}\right)$.

Combining hydrostatic balance and the ideal gas law, we can write:

$$
\frac{\partial p}{\partial z}=-\frac{g p}{R T}
$$

whence

$$
\frac{\partial z}{\partial p}=-\frac{R T}{g p}
$$

The height of a given pressure surface is dependent on the surface pressure $p_{s}$ and the average temperature below that pressure surface:

$$
z(p)=\frac{R}{g} \int_{p}^{p_{s}} T \frac{d p}{p}=\frac{R}{g} \int_{p}^{p_{s}} T d \ln p=\frac{R}{g}\langle T\rangle \int_{p}^{p_{s}} d \ln p
$$

using the definition of $\langle T\rangle$ given in the question.
Thus $\Delta z=z(500 \mathrm{hPa})-z(1000 \mathrm{hPa})$ is given by:

$$
\Delta z=\frac{R\langle T\rangle}{g} \ln 2
$$

a. Compute the thickness of the surface to 500 mbar layer at $30^{\circ}$ and $60^{\circ}$ latitude assuming that the surface temperatures computed in Q2b. above extend uniformly up to 500 mbar .
$\Delta z=\frac{287 \times 310.75}{9.81} \ln 2=6301.6$ at $30^{\circ}$
$\Delta z=\frac{270.87 \times 310.75}{9.81} \ln 2=5947.4$ at $60^{\circ}$
b. Figs.7.4 and 7.25 (of Chapter 7) show 500 mbar and surface pressure analyses for 12GMT on June 21, 2003. Calculate $\langle T\rangle$ for the 1000 mbar to 500 mbar layer at the center of the 500 mbar trough at
$50^{\circ} \mathrm{N}, 120^{\circ} \mathrm{W}$ and at the center of the ridge at $40^{\circ} \mathrm{N}, 90^{\circ} \mathrm{W}$. [N.B. You will need to convert from surface pressure, $p_{s}$, to height of the 1000 hPa surface, $z_{1000}$; to do so use the (approximate) formula

$$
z_{1000} \cong 10\left(p_{s}-1000\right)
$$

where $z_{1000}$ is in meters and $p_{s}$ is in $h \mathrm{~Pa}$.] Is $\langle T\rangle$ greater in the ridge or the trough? Comment on and physically interpret your result.
We know that:

$$
\langle T\rangle=\frac{g \Delta z}{R \ln 2}
$$

In the trough: $p_{s}=1008 \mathrm{~h} \mathrm{~Pa}$ and so $z_{1000} \cong 10(1008-1000)=80 \mathrm{~m}$; $z_{500}=5520 \mathrm{~m}$.
Thus $\Delta z=5520-80=5440 \mathrm{~m}$ and so $\langle T\rangle=\frac{9.81 \times 5440}{287 \times \ln 2} K=268 \mathrm{~K}$.
In the ridge: $p_{s}=1016 \mathrm{hPa}$ and so $z_{1000} \cong 10(1016-1000)=160 \mathrm{~m}$;
$z_{500}=5820 \mathrm{~m}$
Thus $\Delta z=5820-160=5660 \mathrm{~m}$ and so $\langle T\rangle=\frac{9.81 \times 5660}{287 \times \ln 2} K=279 \mathrm{~K}$.
The ridge, then, is some 11 K warmer than the trough. This is reasonable since in the ridge (trough) air has pushed from the south (north) and is thus warm (cold).
4. Use the expression for saturated specific humidity, Eq.(4.24) and the empirical relation for saturated vapor pressure $e_{s}(T)$, Eq.(1.4) (where $A=6.11$ mbar and $\beta=0.067{ }^{\circ} \mathrm{C}^{-1}$ and $T$ is in ${ }^{\circ} \mathrm{C}$ ) to compute from the graph of $T(p)$ in the tropical belt shown in Fig.4.9, vertical profiles of saturated specific humidity, $q^{*}(p)$. You will need to look up values of $R$ and $R_{v}$ from Chapter 1.
Compare your $q^{*}$ profiles with observed profiles of $q$ in the tropics shown in Fig.5.15. Comment?
We evaluate:
$q_{*}=\left(\frac{R}{R_{v}}\right) \frac{e_{s}(T)}{p}$ where $R_{v}=461.39 \mathrm{~J} \mathrm{~kg}^{-1} \mathrm{~K}^{-1} ; R=287.05 \mathrm{~J} \mathrm{~kg}^{-1} \mathrm{~K}^{-1}$ with $e_{s}=A \exp (\beta T)$ where $A=6.11$ mbar and $\beta=0.067{ }^{\circ} \mathrm{C}^{-1}$.
Thus, evaluating the various numerical factors, we have:
$q_{*}=0.62214 \times \frac{6.11 \exp \left(0.067 \times T^{\circ} \mathrm{C}\right)}{p(\mathrm{mbar})}$

From observations of $T(p)$ at the equator, we evaluate $q_{*}$ from the above equation and compare to observations of $q$ in the following table.

| $p$ | $T$ (equator) in ${ }^{\circ} \mathrm{C}$ | $q_{*}(\mathrm{~g} / \mathrm{kg})$ | $q(\mathrm{~g} / \mathrm{kg})$ |
| :---: | :---: | :---: | :---: |
| 1000 | 25.6508 | 21.2 | 16.49 |
| 950 | 22.6972 | 18.3 | 14.56 |
| 900 | 19.9835 | 16.2 | 12.52 |
| 850 | 17.6248 | 14.6 | 10.67 |
| 700 | 9.33311 | 10.1 | 6.04 |
| 500 | -5.59631 | 5.2 | 2.26 |
| 400 | -15.9016 | 3.3 | 1.00 |
| 300 | -31.2027 | 1.56 | 0.31 |
| 200 |  |  |  |
| 100 |  |  |  |

Note that at each level in the atmosphere, $q \lesssim q_{*}$ : the air is less than, but reasonably close to saturation at each level.

