

Middle Atmosphere Dynamics

- 12.1.** Suppose that temperature increases linearly with height in the layer between 20 and 50 km at a rate of 2 K km^{-1} . If the temperature is 200 K at 20 km, find the value of the scale height H for which the log-pressure height z coincides with actual height at 50 km. (Assume that z coincides with the actual height at 20 km and let g be a constant.)

Solution: Recalling that $d\Phi = gdz$, we get integrating equation (10.3):

$$\int_{z_0}^{z_T} \frac{gdz}{T} = \left(\frac{R}{H}\right) \int_{z_0}^{z_T^*} Dz^*, \text{ where } z_0 = z_0^* = 20 \text{ km and } T = T_0 + \gamma(z - z_0) \text{ with } \gamma = 2 \times 10^{-3} \text{ K m}^{-1}. \text{ Thus, } \frac{R}{H} (z_T^* - z_0^*) = \frac{g}{\gamma} \ln \left[\frac{T_0 + \gamma(z_T - z_0)}{T_0} \right], \text{ so to have } z_T^* - z_0^* = 30 \text{ km, we require } H = \frac{R(z_T^* - z_0^*)\gamma}{g \ln \left[\frac{T_0 + \gamma(z_T - z_0)}{T_0} \right]} = \frac{287 \times 30 \times 2}{9.8 \times \ln \left(\frac{260}{200} \right)} = 6697 \text{ m. (This scale height corresponds to a mean temperature of } T = 229 \text{ K.)}$$

- 12.2.** Find the Rossby critical velocities for zonal wave numbers 1, 2, and 3 (i.e., for 1, 2, and 3 wavelengths around a latitude circle). Let the motion be referred to a β -plane centered at 45°N , scale height $H = 7 \text{ km}$, buoyancy frequency $N = 2 \times 10^{-2} \text{ s}^{-1}$, and infinite meridional scale ($l = 0$).

Solution: From (12.16) $U_c = \beta [k^2 + f_0^2 / (4N^2H^2)]^{-1}$ where $l = 0$. But here $k = s / (a \cos \phi)$ and $\beta = 2\Omega \cos \phi / a$. Thus, $U_c = 1.619 \times 10^{-11} [s^2 (4.93 \times 10^{-14}) + 1.356 \times 10^{-13}]^{-1}$. Then,

| s | U_c (m/s) |
|-----|-------------|
| 1 | 87.5 |
| 2 | 48.7 |
| 3 | 28.0 |

- 12.3.** Suppose that a stationary linear Rossby wave is forced by flow over sinusoidal topography with height $h(x) = h_0 \cos(kx)$, where h_0 is a constant and k is the zonal wavenumber. Show that the lower boundary condition on the streamfunction ψ can be expressed in this case as

$$(\partial\psi / \partial z) = -hN^2 / f_0.$$

Using this boundary condition and an appropriate upper boundary condition, solve for $\psi(x, z)$ in the case $|m| \gg (1/2H)$ using the equations of Section 12.3.1. How does the position of the trough relative to the mountain ridge depend on the sign of m^2 ?

Solution: Now for topographic forcing $w' = \bar{u} \partial h / \partial x$ at $z = 0$. From the linearized form of the adiabatic thermodynamic energy equation [see (8.46)], we then have at $z = 0$:

$\left(\frac{\partial}{\partial t} + \bar{u} \frac{\partial}{\partial x}\right) \frac{\partial \Phi'}{\partial z} = -w' N^2 = -N^2 \bar{u} \frac{\partial h}{\partial x}$. But $\Phi' = f_0 \psi'$, so the lower boundary condition can be expressed for stationary waves ($\partial / \partial t = 0$) as:

$$\frac{\partial \psi'}{\partial z} = -\frac{N^2 h}{f_0} = -\frac{N^2 h_0}{f_0} \cos(kx). \text{ But from (12.13),}$$

$\psi'(x, z) = \text{Re}\{A \exp[i(kx + mz) + z/2H]\}$ where A is a complex constant. And m is given by (12.15). The lower boundary condition then permits determination of A :

$$\frac{\partial \psi'}{\partial z} = \text{Re}\left\{\left(im + \frac{1}{2H}\right) A \exp(ikx)\right\} = -\frac{N^2 h_0}{f_0} \cos(kx), \text{ at } z=0. \text{ Thus, for } m^2 > 0 \text{ and } |m| \gg 1/2H, \psi'(x, z) = -\frac{N^2 h_0}{mf_0} \sin(kx + mz) \exp\left(\frac{z}{2H}\right).$$

$$\text{For } m^2 < 0 \text{ and } |m| \gg 1/2H, \psi'(x, z) = \frac{N^2 h_0}{|m|f_0} \cos(kx) \exp\left(\frac{z}{2H} - |m|z\right)$$

Thus, in the former case the wave propagates vertically and there is a lee-side trough, while in the latter the geopotential ridge coincides with the mountain ridge and the wave energy decays with height.

- 12.4.** Consider a very simple model of a steady state mean meridional circulation for zonally symmetric flow in a midlatitude channel bounded by walls at $y = 0, \pi/l$ and $z = 0, \pi/m$. We assume that the zonal mean zonal flow \bar{u} is in thermal wind balance and that the eddy momentum and heat fluxes vanish. For simplicity we let $\rho_0 = 1$ (Boussinesq approximation), and let the zonal force owing to small-scale motions be represented by a linear drag: $\bar{X} = -\gamma\bar{u}$. We assume that the diabatic heating has the form $J/c_p = (H/R) J_0 \cos ly \sin mz$, and we let N and f be constants. Eqs. (12.1), (12.2), (12.3), and (12.4) then yield the following:

$$-f_0 \bar{v}^* = -\gamma \bar{u} \quad (\text{i})$$

$$+N^2 H R^{-1} \bar{w}^* = +\bar{J}/c_p \quad (\text{ii})$$

$$\bar{v}^* = -\frac{\partial \bar{X}^*}{\partial z}; \quad \bar{w}^* = \frac{\partial \bar{X}^*}{\partial y} \quad (\text{iii}, \text{b})$$

$$f_0 \bar{u} / \partial z + R H^{-1} \partial \bar{T} / \partial y = 0 \quad (\text{iv})$$

Assuming that there is no flow through the walls, solve for the residual circulation defined by \bar{X}^* , \bar{v}^* , and \bar{w}^* .

Solution: From (ii), we immediately get $\bar{w}^* = (J_0/N^2) \cos(ly) \sin(mz)$. Then from (iii), $\bar{X}^* = (J_0 N^{-2} l^{-1}) \sin(ly) \sin(mz)$ and $\bar{v}^* = -(J_0/N^2) (m/l) \sin(ly) \cos(mz)$.

- 12.5.** For the situation of problem 12.4 solve for the steady state zonal wind and temperature fields \bar{u} and \bar{T} .

Solution: Now from equation (i) of problem 12.4 we find that $\bar{u} = \left(\frac{f_0}{\gamma}\right) \bar{v}^* = -\left(\frac{f_0}{\gamma}\right) \left(\frac{J_0}{N^2}\right) \left(\frac{m}{l}\right) \sin(ly) \cos(mz)$. And application of the thermal wind equation then gives $\bar{T} = \left(\frac{f_0^2}{\gamma}\right) \left(\frac{H}{R}\right) \left(\frac{J_0}{N^2}\right) \left(\frac{m}{l}\right)^2 \cos(ly) \sin(mz)$.

- 12.6.** Find the geopotential and vertical velocity fluctuations for a Kelvin wave of zonal wave number 1, phase speed 40 m s^{-1} , and zonal velocity perturbation amplitude 5 m s^{-1} . Let $N^2 = 4 \times 10^{-4} \text{ s}^{-2}$.

Solution: From (12.38) $\Phi' = u'(v/k) = u'c$, so $|\Phi'| = (40)(5) = 200 \text{ m}^2 \text{ s}^{-2}$. From (12.37) $w' \approx (vm/N^2)\Phi'$ since $|m| \gg 1/(2H)$. Also, $v = ck = cs/a$, where s is the planetary wavenumber. Now, $m^2 \approx (N/c)^2 = 4 \times 10^{-4}/1600 = 25 \times 10^{-8}$, so that $|m| = 5 \times 10^{-4} \text{ m}^{-1}$. Substituting into the formula for vertical velocity:

$$|w'| = (40) \left(6.37 \times 10^6\right)^{-1} \left(5 \times 10^{-4}\right) (200) \left(4 \times 10^{-4}\right)^{-1} = 1.57 \times 10^{-3} \text{ m s}^{-1}$$

- 12.7.** For the situation of Problem 12.6 compute the vertical momentum flux $M \equiv \rho_0 \overline{u'w'}$. Show that M is constant with height.

Solution: Now by definition: $\rho_0 = \rho_s \exp(-z^*/2H)$. Thus,

$M = \rho_s \exp(-z^*/H) \overline{u'w'}$. But $u' = U_0 \cos(kx + mz - vt) \exp(+z^*/2H)$, and from Problem 12.6, $w' = (c^2 km/N^2) U_0 \cos(kx + mz - vt) \exp(+z^*/2H)$. Thus, for $\rho_s = 1 \text{ kg m}^{-3}$

$$\begin{aligned} M &= \rho_s \left(\frac{U_0^2 c^2 km}{N^2} \right) \overline{\cos^2(kx + mz - vt)} = \frac{(5)^2 (40)^2 (5 \times 10^{-4}) (0.5)}{(6.37 \times 10^6) (4 \times 10^{-4})} \\ &= 3.92 \times 10^{-3} \text{ kg m}^{-1} \text{ s}^{-1} \end{aligned}$$

- 12.8.** Determine the form for the vertical velocity perturbation for the Rossby-gravity wave corresponding to the u' , v' , and Φ' perturbations given in (12.44).

Solution: From (12.37) and (12.44), and with $|m| \gg 1/(2H)$:

$$\hat{w} \approx -\left(vm/N^2\right) \hat{\Phi} = -\hat{v} \left(\frac{iv^2 m}{N^2}\right) y \exp\left[-\frac{\beta |m| y^2}{2N}\right]$$

- 12.9.** For a Rossby-gravity wave of zonal wave number 4 and phase speed -20 m s^{-1} , determine the latitude at which the vertical momentum flux $M \equiv \rho_0 \overline{u'w'}$ is a maximum.

Solution: From problem 12.8 and equation (12.44): $\overline{u'w'} = Ay^2 \exp\left[-\frac{\beta|m|y^2}{N}\right]$, where A is a constant. Then $\partial\overline{u'w'}/\partial y = 0 \rightarrow y^2(\beta|m|/N) = 1$, or $y_{\max} = \pm v\beta^{-1}(1 + kv/\beta)^{-1/2}$, where $v = ck = cs/a = (-20)(4)(6.37 \times 10^6)^{-1} \text{ s}^{-1}$. Thus, $y_{\max} = \pm \frac{cs}{2\Omega(1 + 0.5sv/\Omega)^{1/2}} = \frac{\pm 20 \times 4}{(1.458 \times 10^{-4})(0.664)^{1/2}} = \pm 677.7 \text{ km}$, or about 6° latitude north and south of the equator.

- 12.10.** Suppose that the mean zonal wind shear in the descending westerlies of the equatorial QBO can be represented analytically on the equatorial β -plane in the form $\partial u/\partial z = \Lambda \exp(-y^2/L^2)$, where $L = 1200 \text{ km}$. Determine the approximate meridional dependence of the corresponding temperature anomaly for $|y| \ll L$.

Solution: From (12.45) $\frac{\partial^2 \bar{T}}{\partial y^2} = -\frac{\beta H}{R} \frac{\partial \bar{u}}{\partial z} = -\frac{\beta H \Lambda}{R} \exp(-y^2/L^2)$. An approximate solution can be obtained by expanding the y -dependence as follows:

$$\bar{T} = T_0 + T_1 \exp(-y^2/L^2) + T_2 (y^2/L^2) \exp(-y^2/L^2) + \text{higher-order terms,}$$

where T_0 , T_1 , and T_2 are constants. Substituting into the above equation and equating terms with the same y dependence gives $\frac{2}{L^2}(T_1 - T_2) = \frac{\beta H \Lambda}{R}$ and $-\frac{6}{L^4}T_2 + \frac{4}{L^4}T_1 - \frac{4}{L^4}T_2 = 0$. Thus, $T_2 = \frac{4}{10}T_1$ and $T_1 = \left(\frac{5}{6}\right) \frac{\beta H \Lambda L^2}{R} \exp(-y^2/L^2)$.

- 12.11.** Estimate the TEM residual vertical velocity in the westerly shear zone of the equatorial QBO assuming that radiative cooling can be approximated by Newtonian cooling with a 20-day relaxation time, the vertical shear is 20 m s^{-1} per 5 km , and the meridional half-width is 12° latitude.

Solution: In this case the temperature tendency term can be neglected, and (12.2) is an approximate balance between adiabatic warming and Newtonian cooling: $(N^2 H/R) \bar{w}^* \approx \bar{J}/c_p = -\alpha(\bar{T} - T_0)$. But from Problem 12.10, near the equator we have $\bar{T} - T_0 \approx +\frac{\beta H \Lambda L^2}{R} \left(\frac{5}{6}\right) \exp(-y^2/L^2)$. Thus, $\bar{w}^* \approx -\frac{\alpha \beta \Lambda L^2}{N^2} \left(\frac{5}{6}\right) \exp(-y^2/L^2)$.

Now, $\alpha = 1/20 \text{ days} = 5.8 \times 10^{-7} \text{ s}^{-1}$. So,

$$\bar{w}^* \approx -\frac{(5/6)(5.8 \times 10^{-7})(2.29 \times 10^{-11})(4 \times 10^{-3})(1.33 \times 10^6)^2}{(4 \times 10^{-4})} = -1.96 \times 10^{-4} \text{ m s}^{-1},$$

or $\bar{w}^* \approx -16.9 \text{ m day}^{-1}$ at the equator. (Note that there is subsidence corresponding to the westerly shear phase of the QBO.)