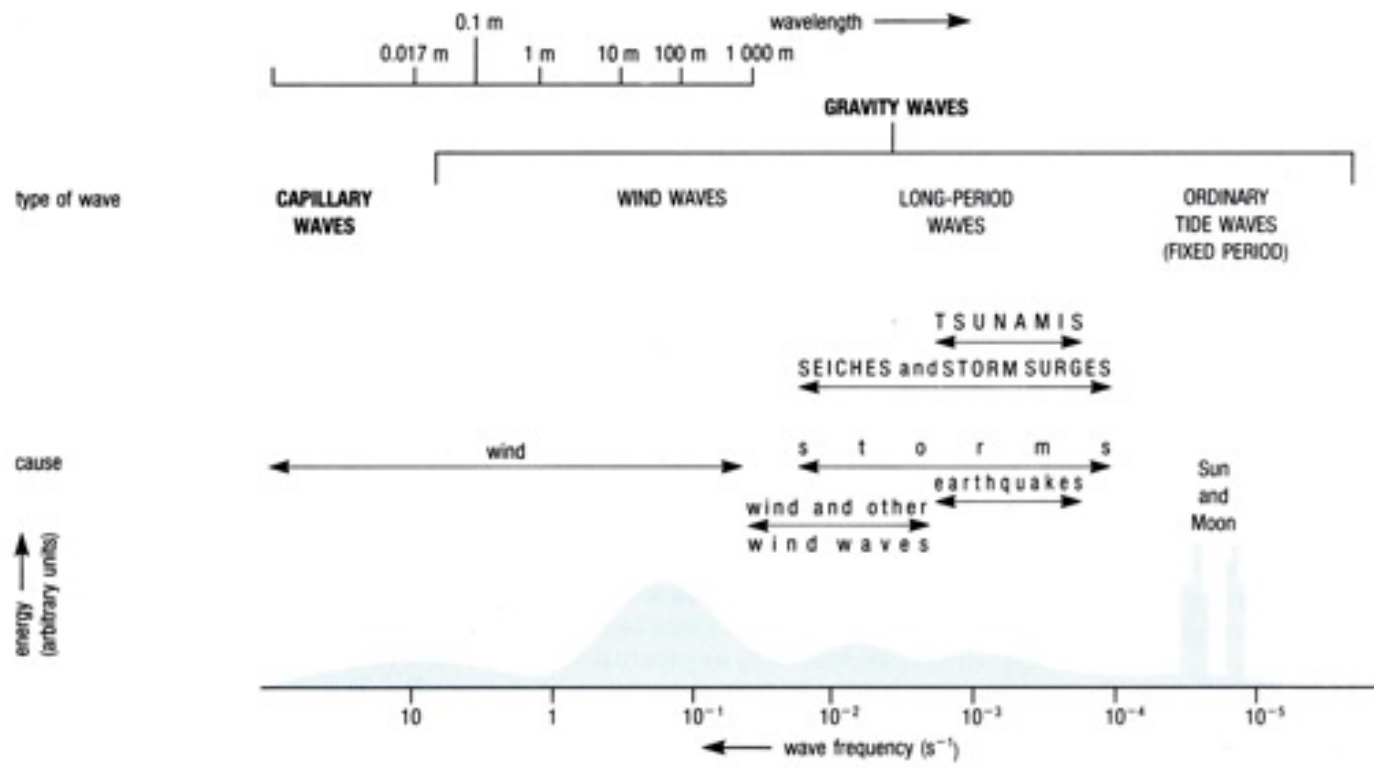


Waves

AOS660, Professor McKinley

14 November 2013

- Wave basics, a review
 - Dispersion relationship
 - Dispersive vs. Non-dispersive
- Examples
 - Shallow and Deep water Waves
 - Surface
 - Internal
 - Tsunami
 - Planetary Waves
 - Kelvin
 - Rossby



What is a wave?

- Waves transfer energy without transfer of mass
- To exist, waves require
 - Inertia – set up by an initial disturbance
 - Restoring force

Particles have orbital trajectory

16

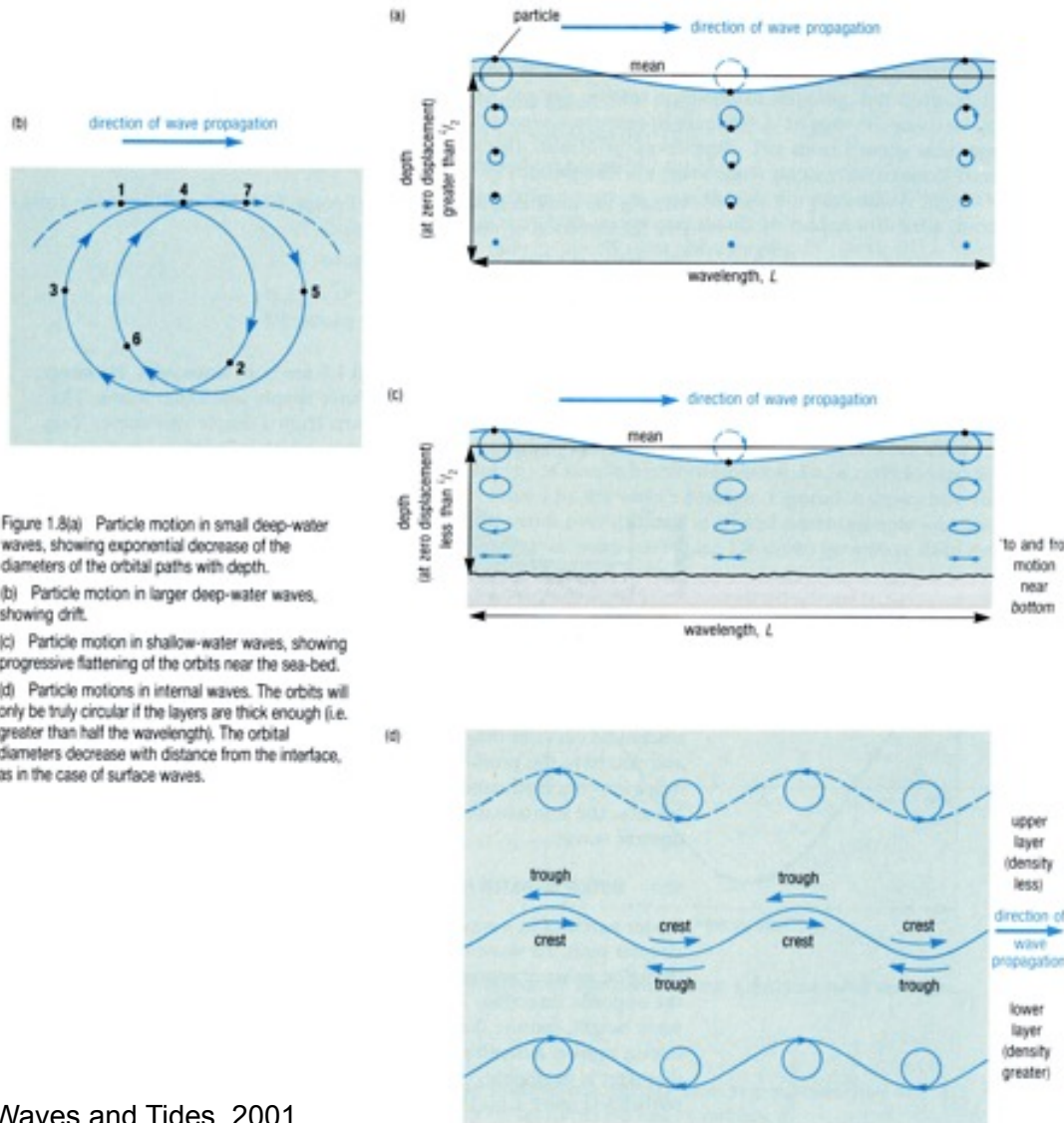


Figure 1.8(a) Particle motion in small deep-water waves, showing exponential decrease of the diameters of the orbital paths with depth.
 (b) Particle motion in larger deep-water waves, showing drift.
 (c) Particle motion in shallow-water waves, showing progressive flattening of the orbits near the sea-bed.
 (d) Particle motions in internal waves. The orbits will only be truly circular if the layers are thick enough (i.e. greater than half the wavelength). The orbital diameters decrease with distance from the interface, as in the case of surface waves.

Wave dimensions

- Space
 - Wavelength (L)
 - Wavenumber (e.g. $k = 2\pi/L$)
 - Amplitude (A)
- Time
 - Period (T) (s)
 - Wave frequency f (1/s) or Angular Frequency σ (rad/s)
- Velocity
 - Phase speed (c), Group velocity (c_g) (m/s)

Dispersion relation

- Relates frequency to wave number
 - Though often given as speed to wave number relationship
- Fundamental property of the wave form!
 - Tells how the restoring force acts

Deriving the shallow water dispersion relationship

- Write down the equation of motion, shallow water
 - Ignore rotation, friction
 - Pressure gradient in terms of surface height
- Apply the perturbation method to consider small variations around the mean state
- Develop a second order DE for the surface height (use continuity)
- Assume surface height has a wave-like solution (e.g. $h' = Ae^{ik(x-ct)}$)
- Plug into the 2nd order DE and solve

Shallow water gravity wave speed, surface, no mean flow

$$\rho_1 = 1000 \text{ kg/m}^3$$

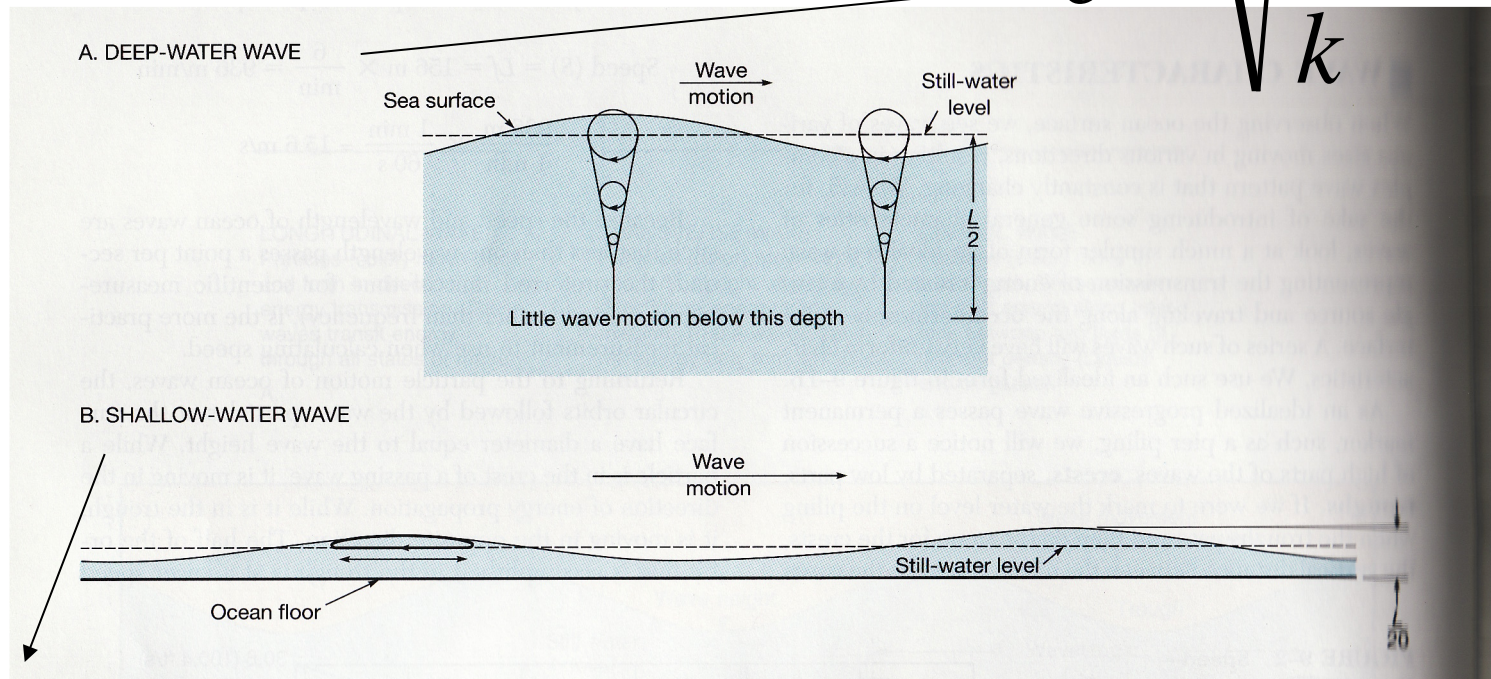
$$\rho_2 = 1 \text{ kg/m}^3$$

$$\rho' = \rho_1 - \rho_2 = 999 \text{ kg/m}^3$$

$$\rho' / \rho_1 \sim 1$$

$$c = \sqrt{gH}$$

$$c = \sqrt{\frac{g}{k}}$$



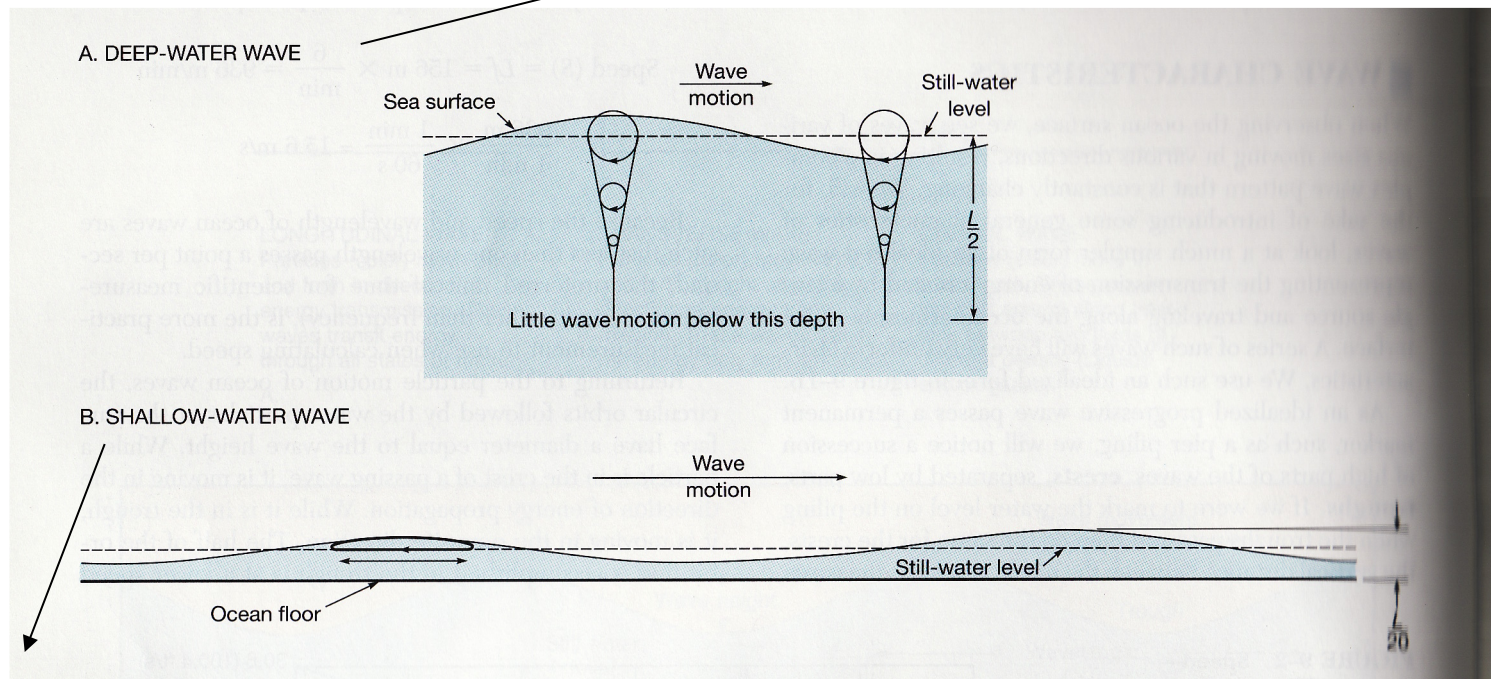
$$c = \sqrt{gH}$$

Dispersive vs. Non-dispersive

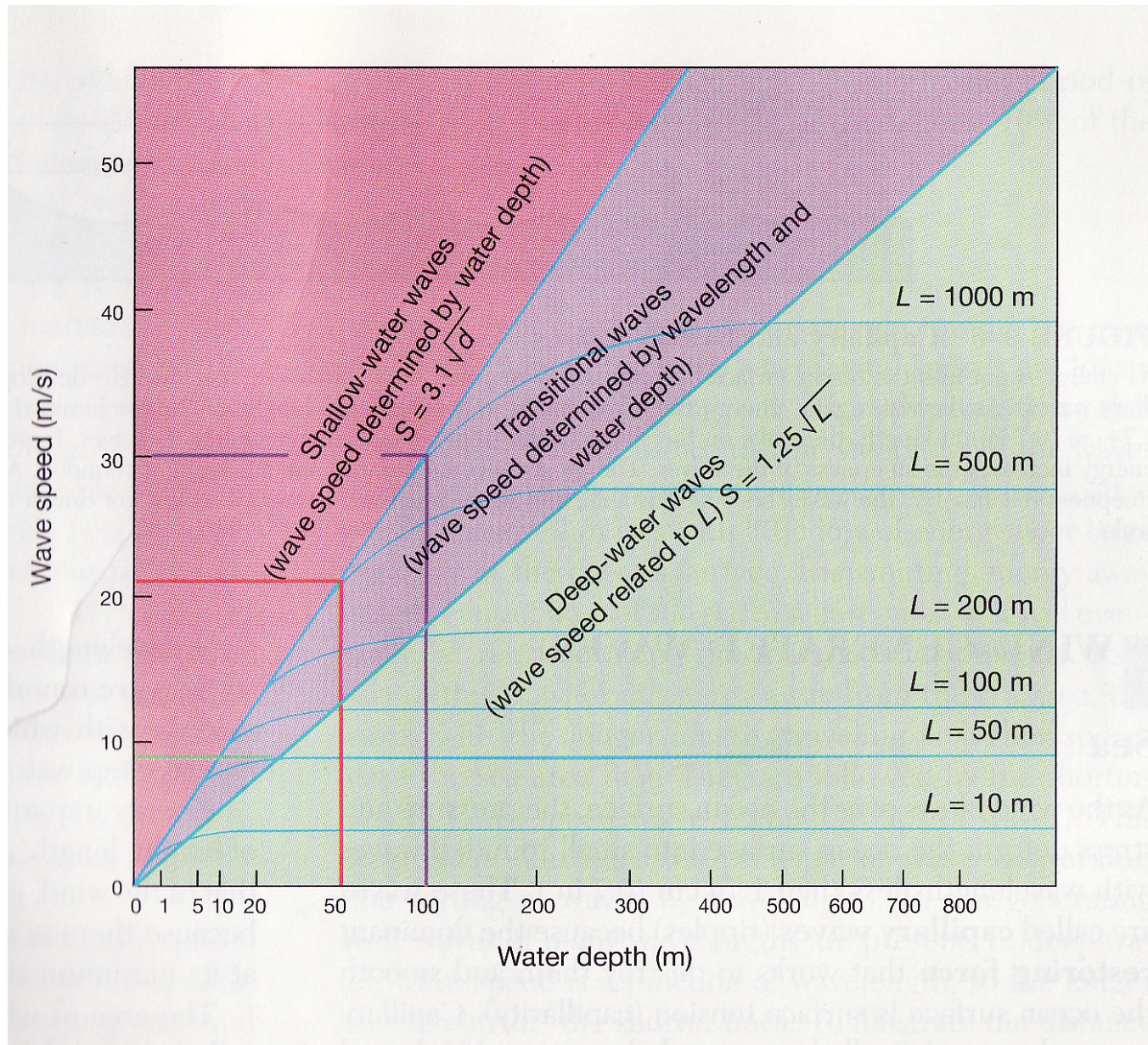
- If the wave speed depends on the wave frequency (or wave length), then the wave speed changes depending on the wave form
- If no such dependence, all waves of this type (no matter their size), propagate at the same speed

Which of these waves is dispersive?

$$c = \sqrt{\frac{g}{k}} = \sqrt{\frac{gL}{2\pi}}$$

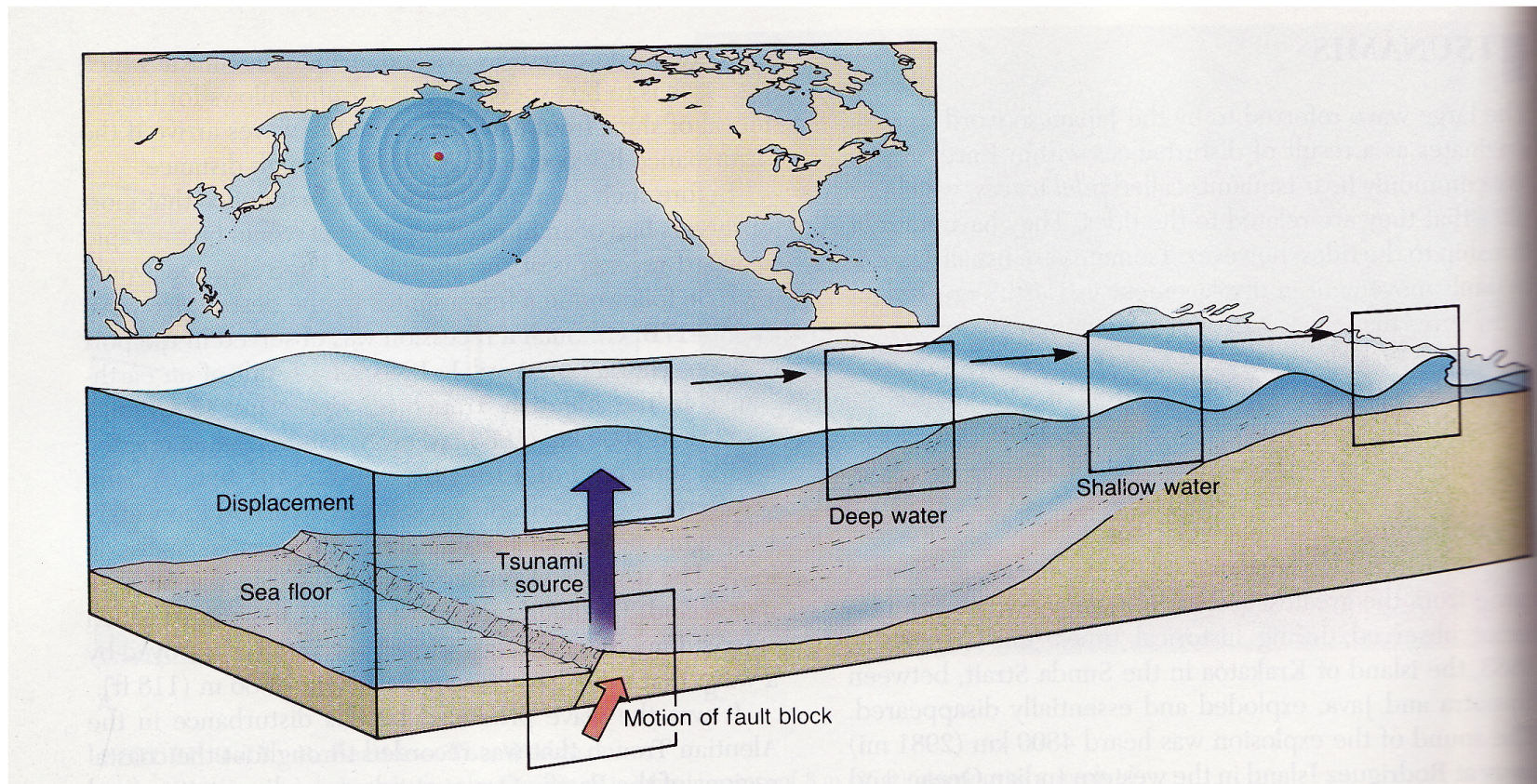


$$c = \sqrt{gH}$$





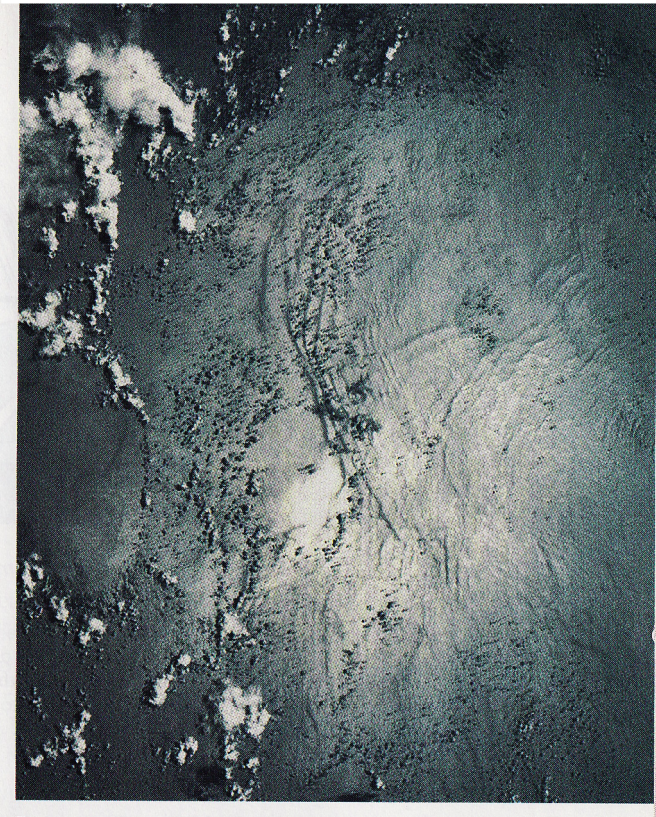
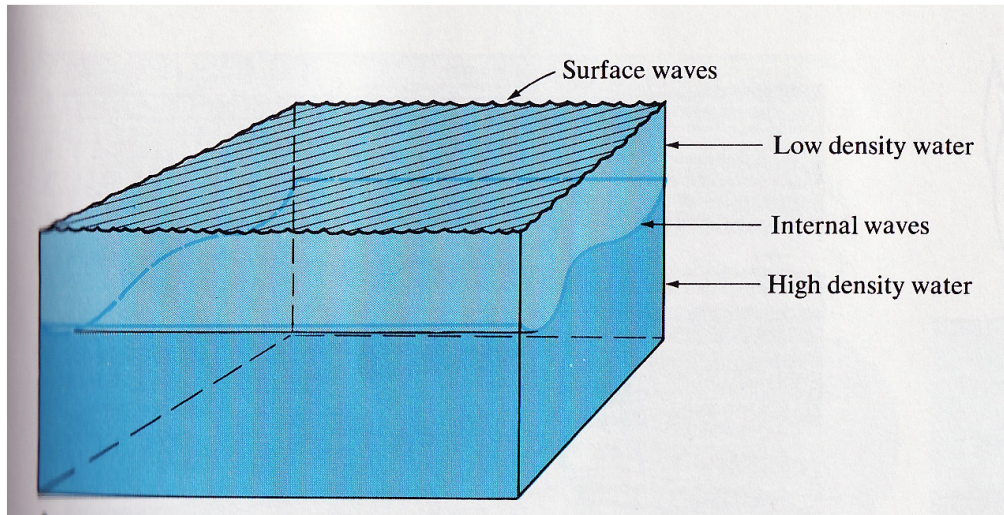
Tsunami



A Tsunami is essentially a
shallow water wave

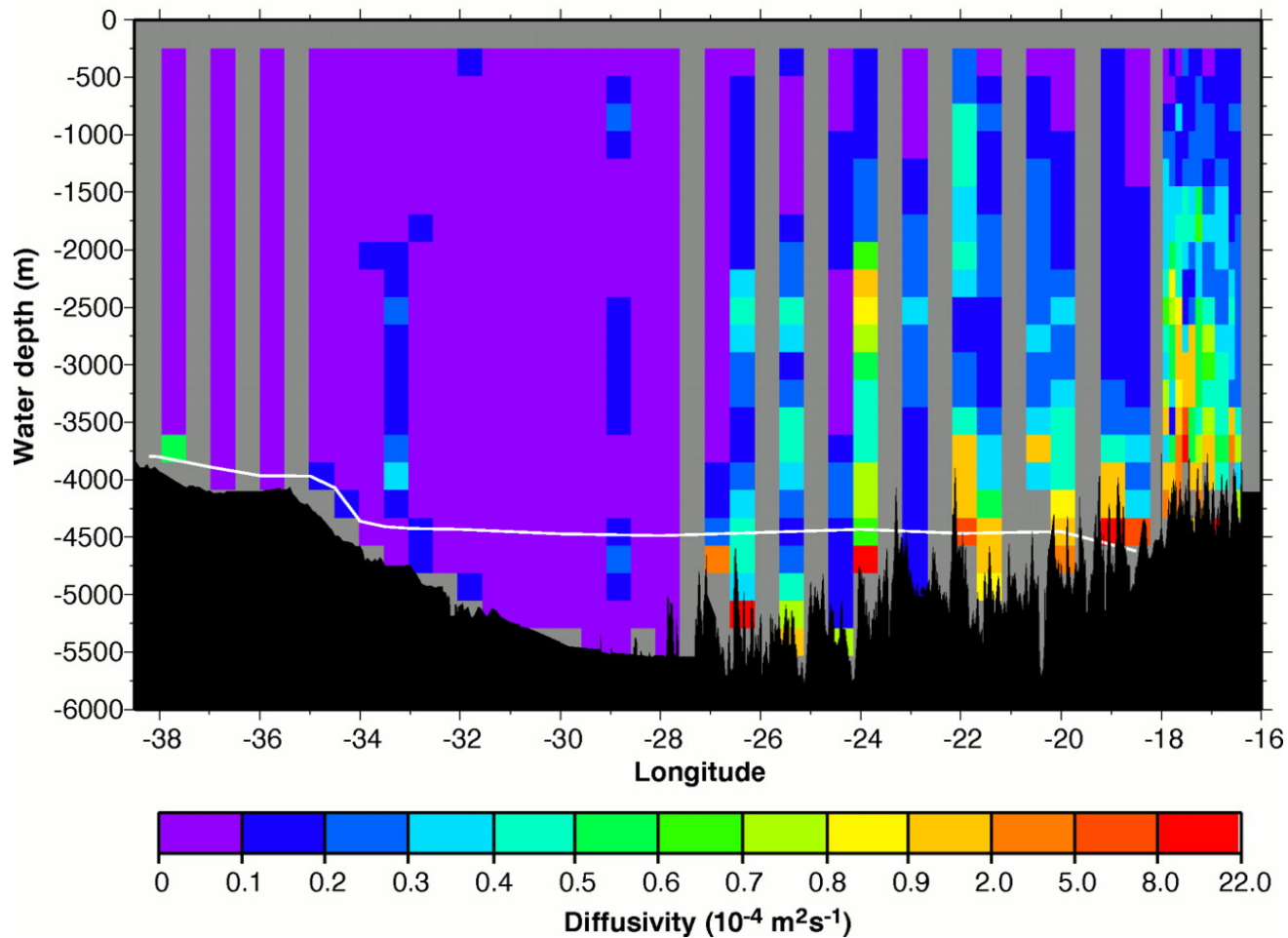
How fast does it propagate?

Internal waves



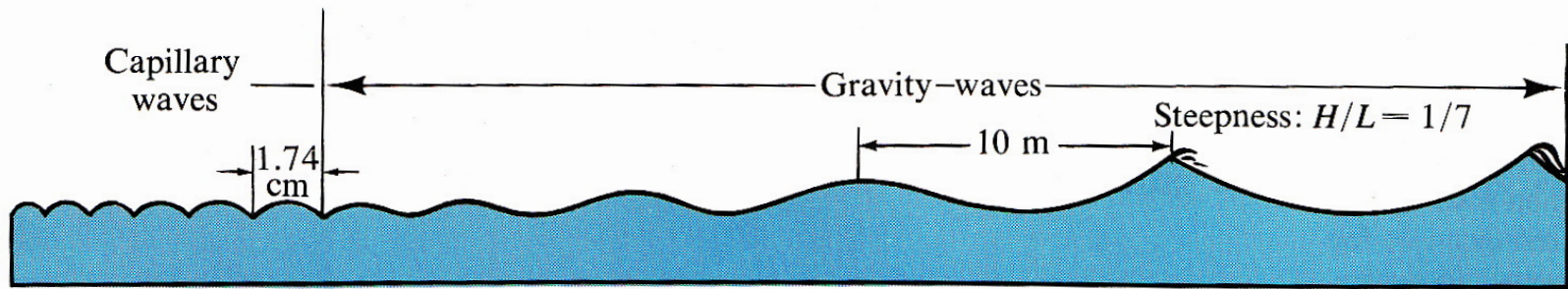
Internal mixing due to waves and wave breaking

Brazil Basin



Polzin et al. 1997

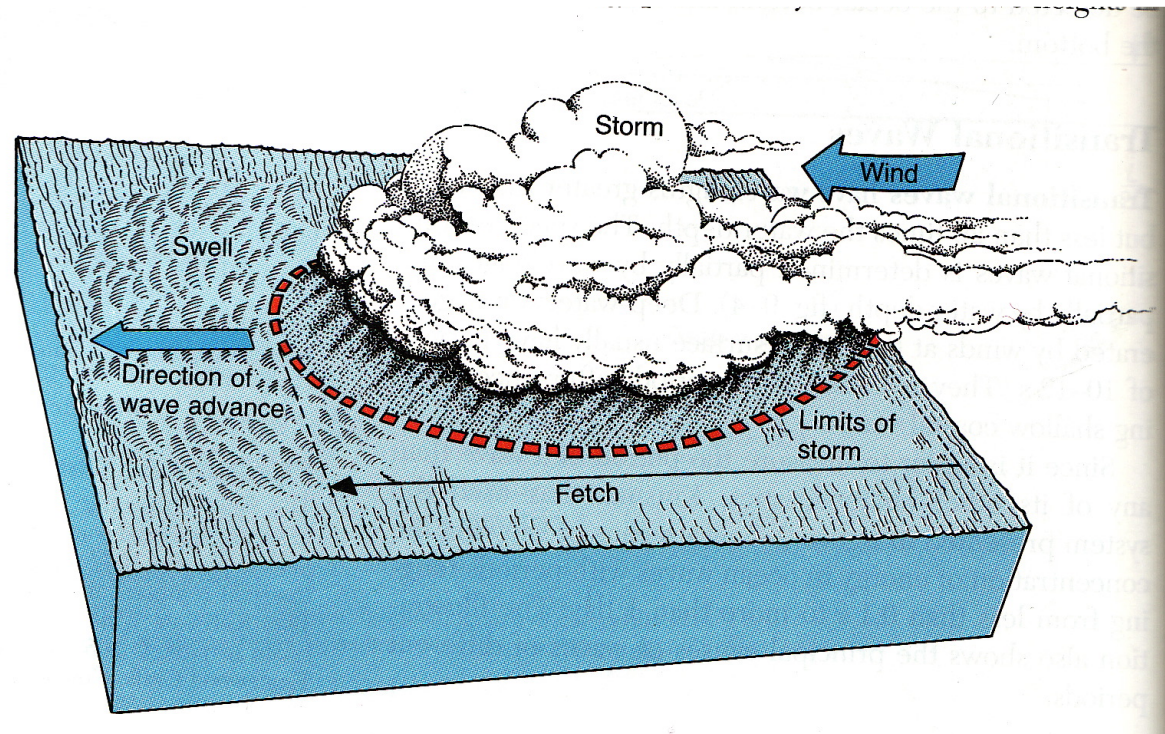
From the sea to the shore

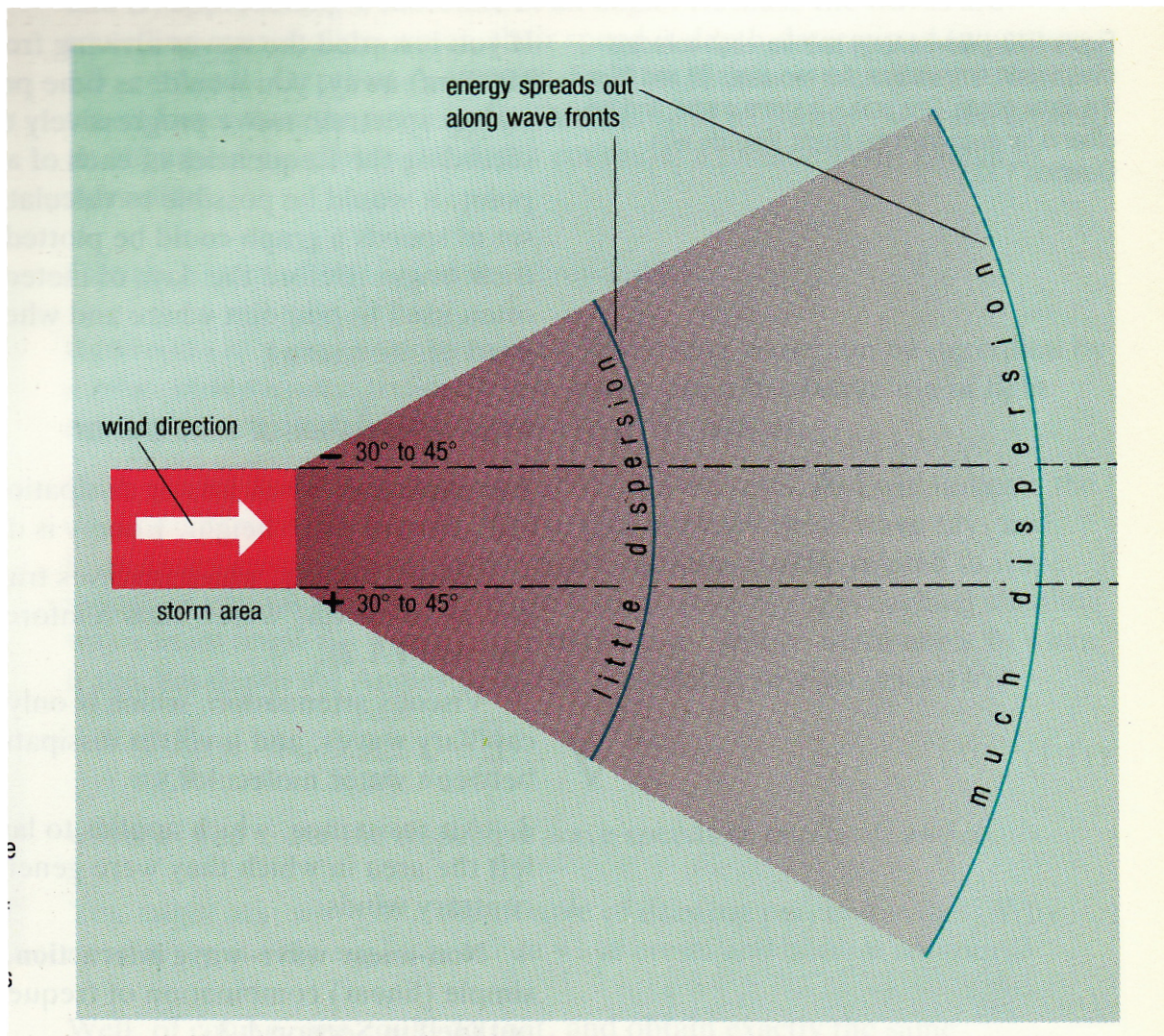


Swell

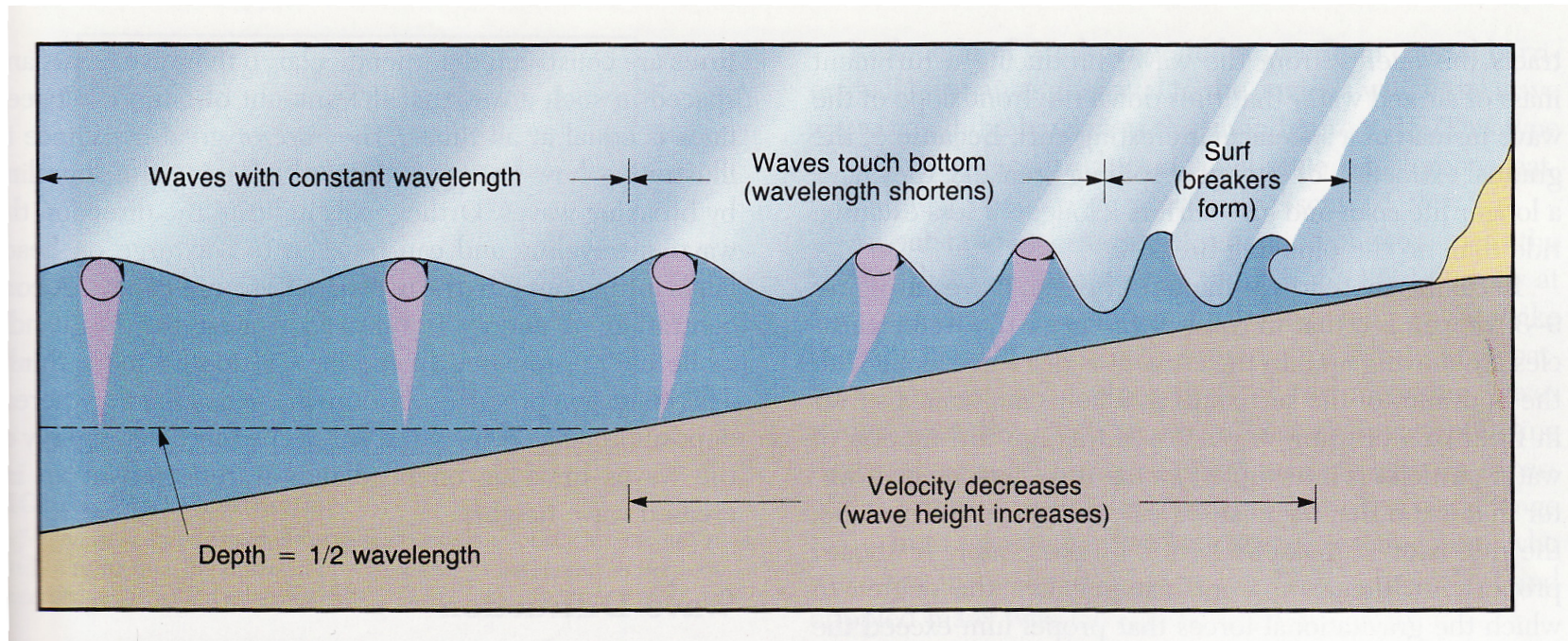
FIGURE 9-7 The Sea and Swell.

As wind blows across the "sea," wave size increases with increased wind speed, fetch, and duration. As waves advance beyond the sea, they continue to advance across the ocean surface as "swell," free waves that are not driven by the wind but sustained by the energy they obtained in the sea.

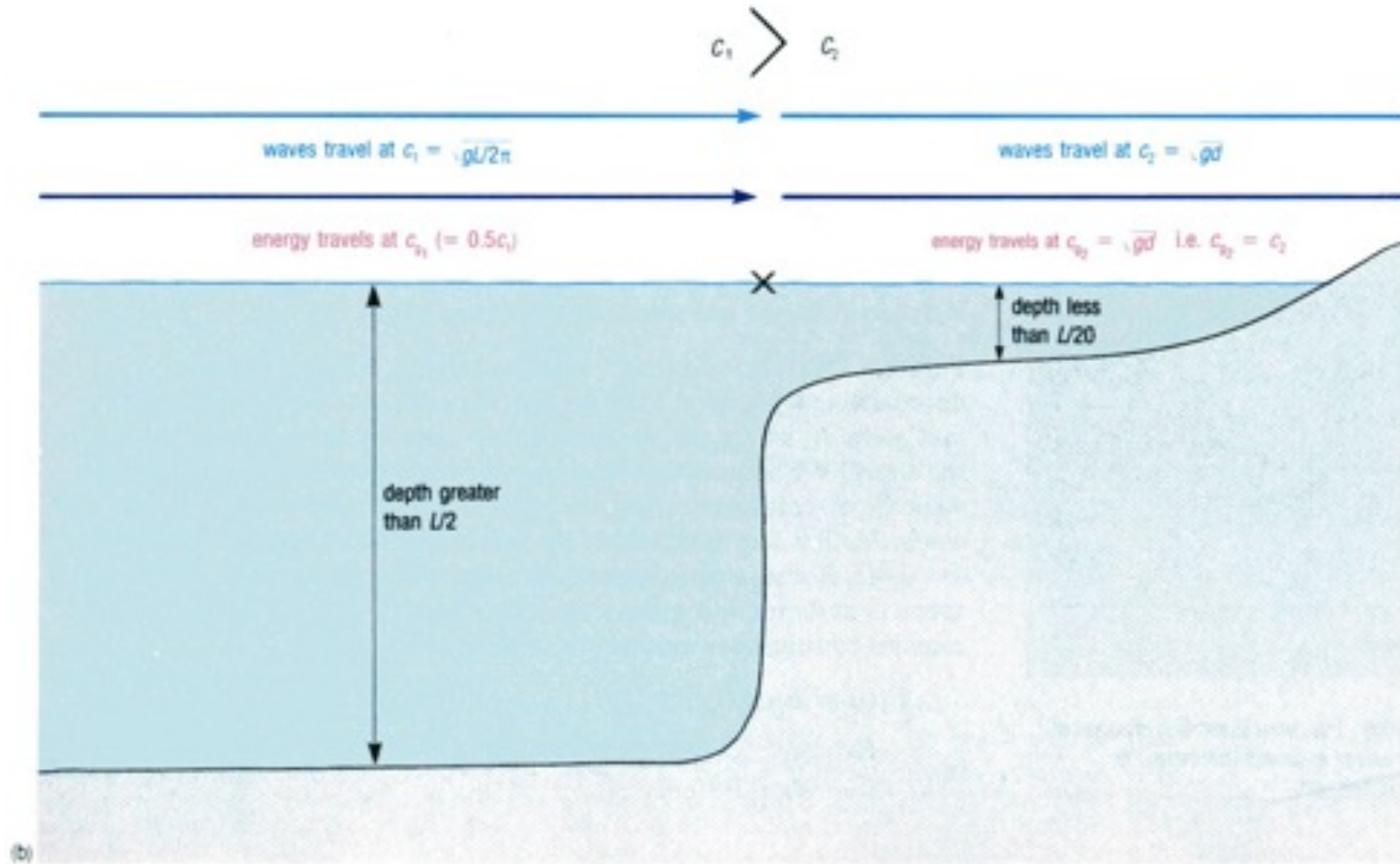


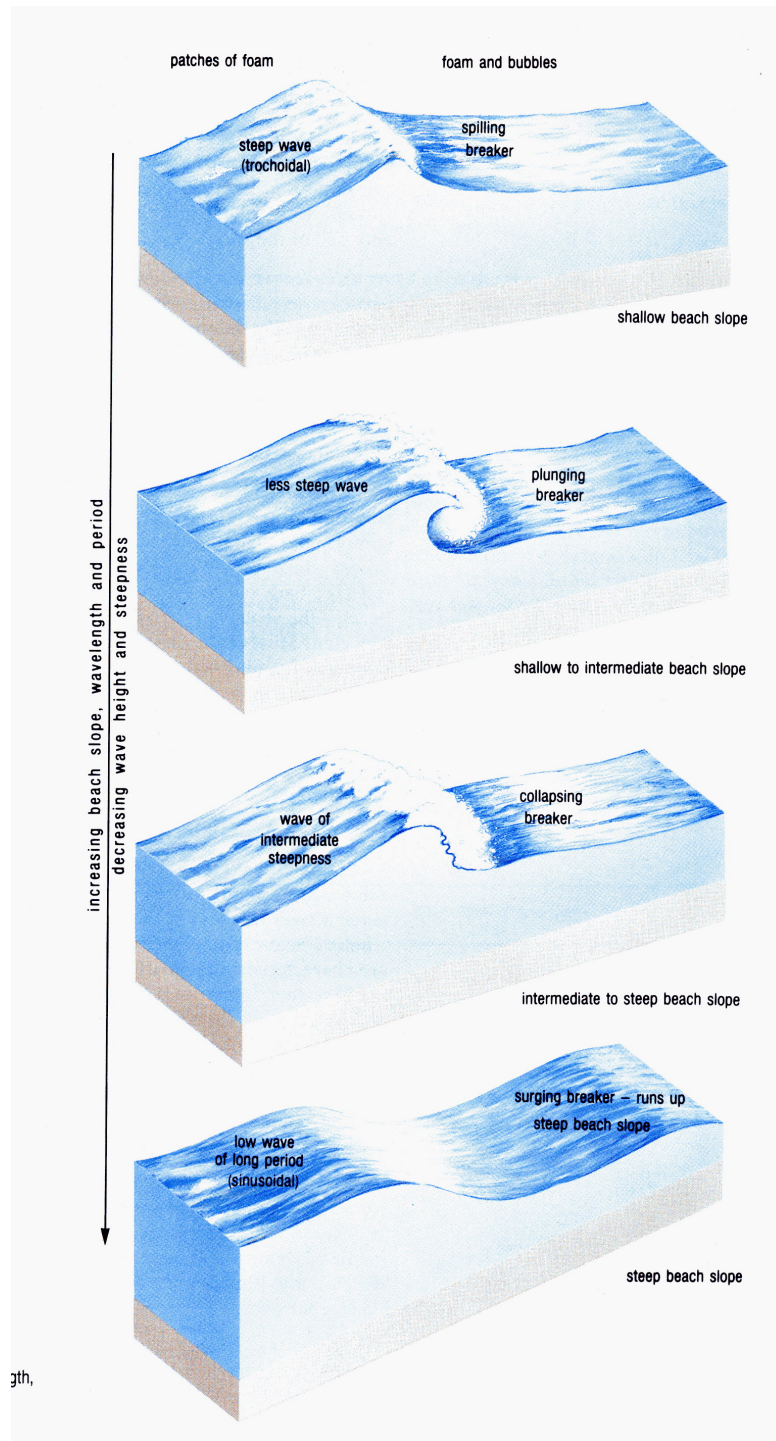


In the surf zone



Energy transformation as wave approaches shore

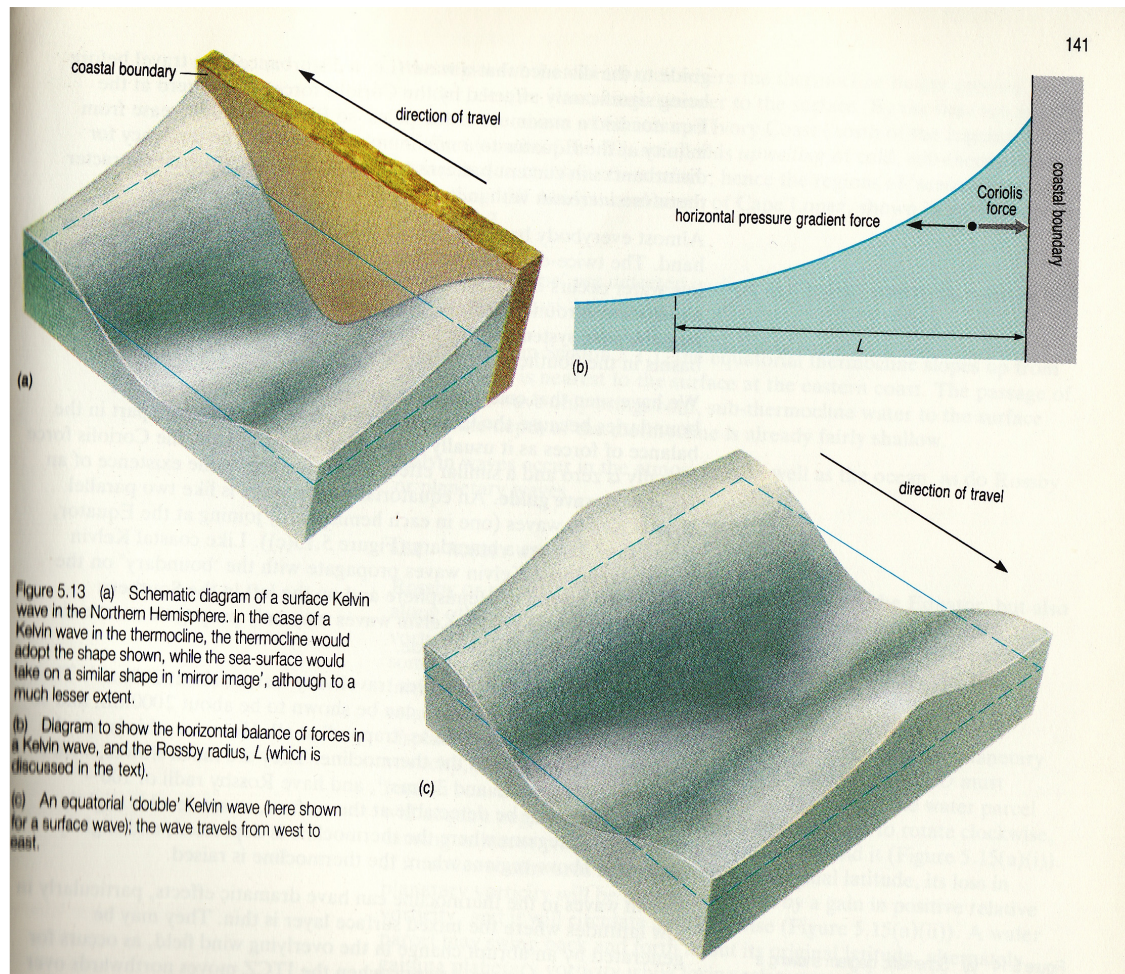




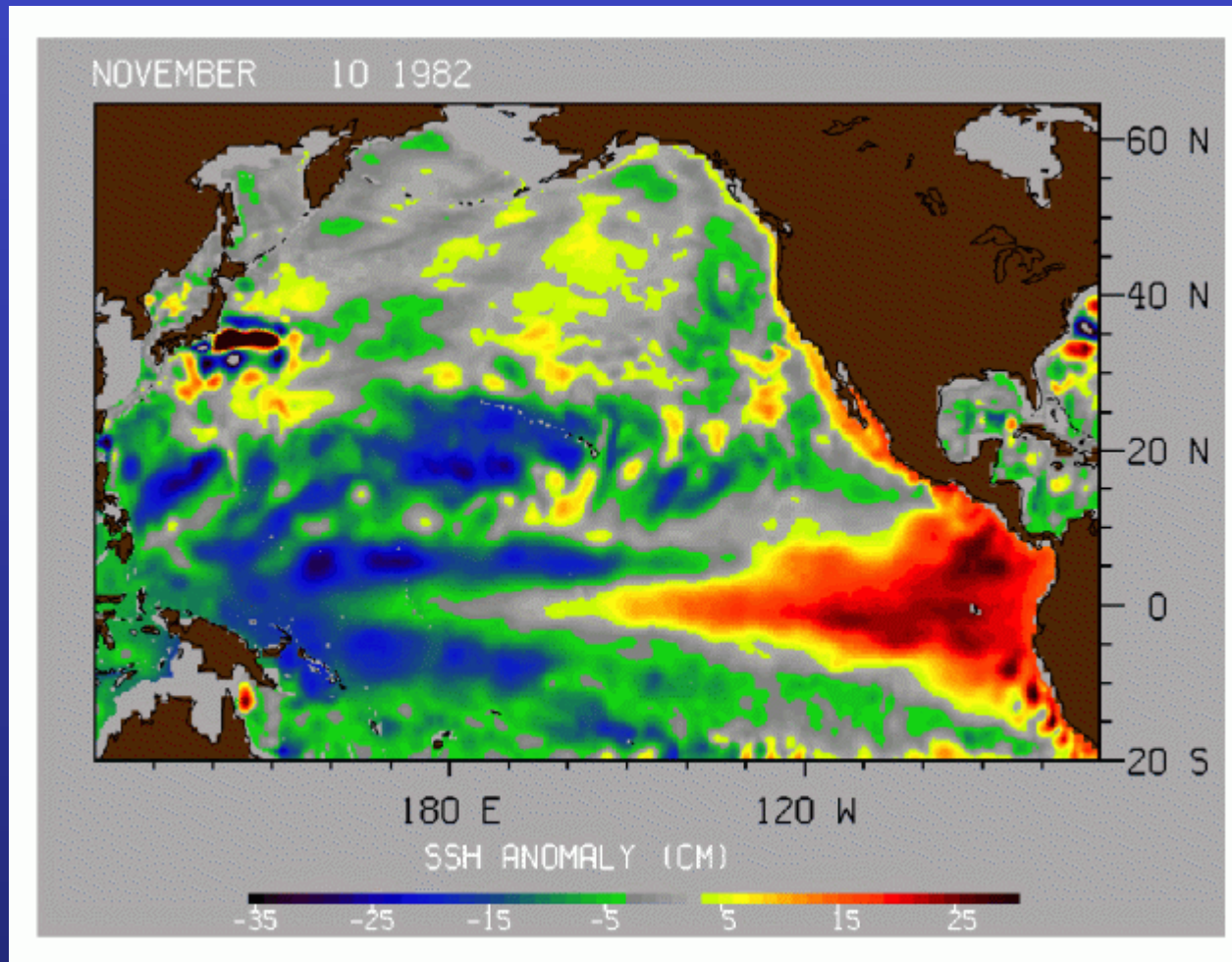
gth,

Planetary Waves

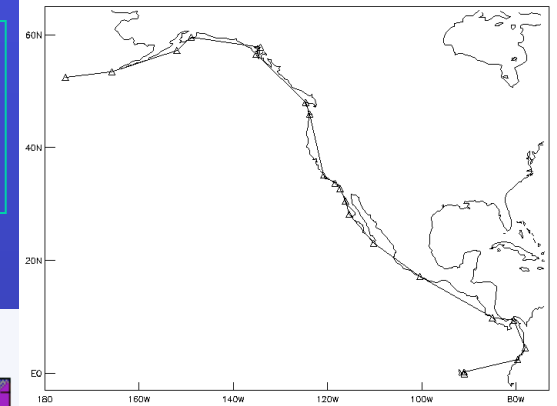
Kelvin wave: A kind of surface gravity wave, shallow water



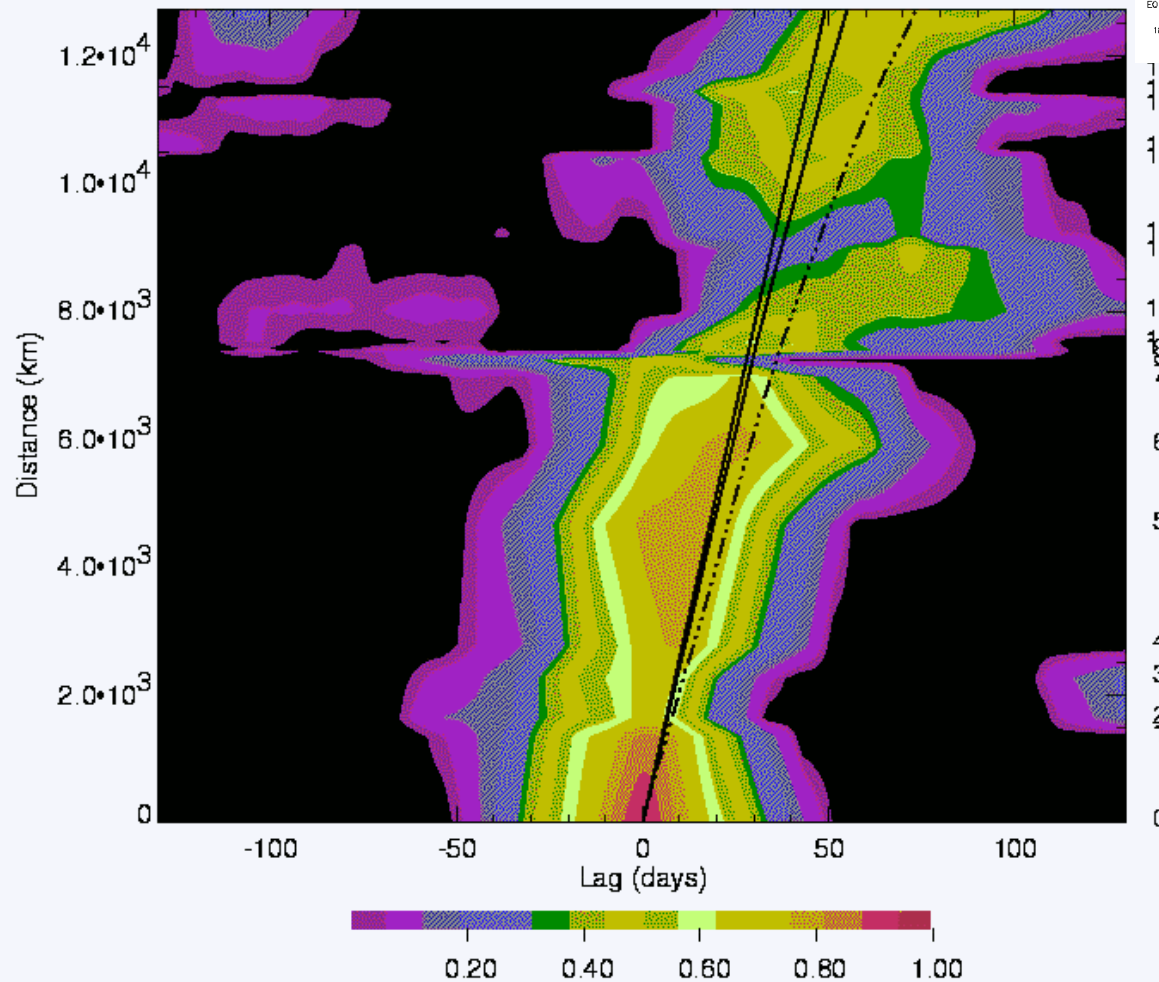
Coastal Kelvin Waves



Coastal Kelvin Waves

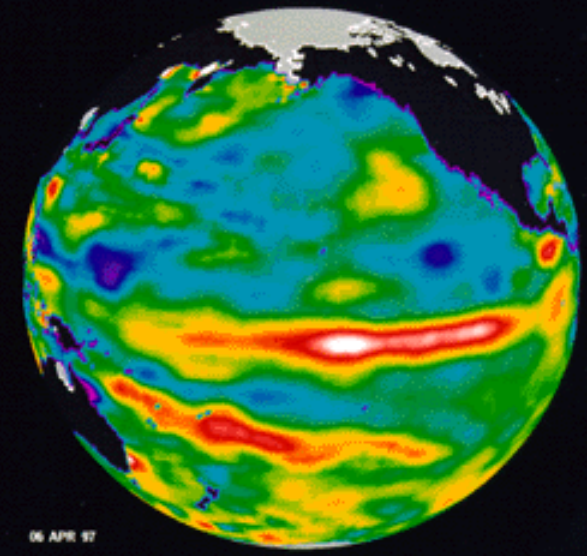
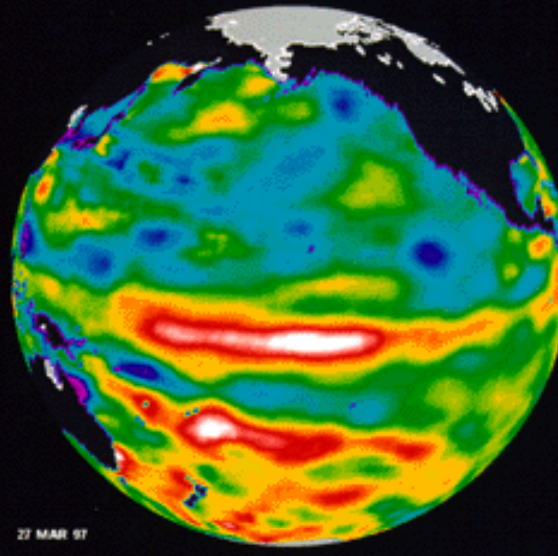
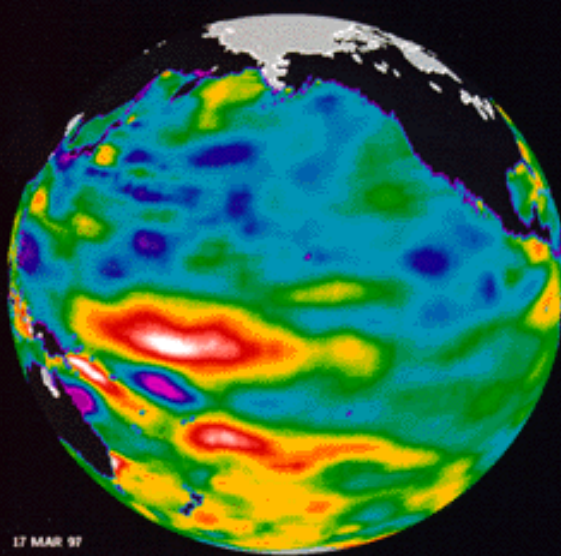


Correlation of 1982-1983 sea level anomalies

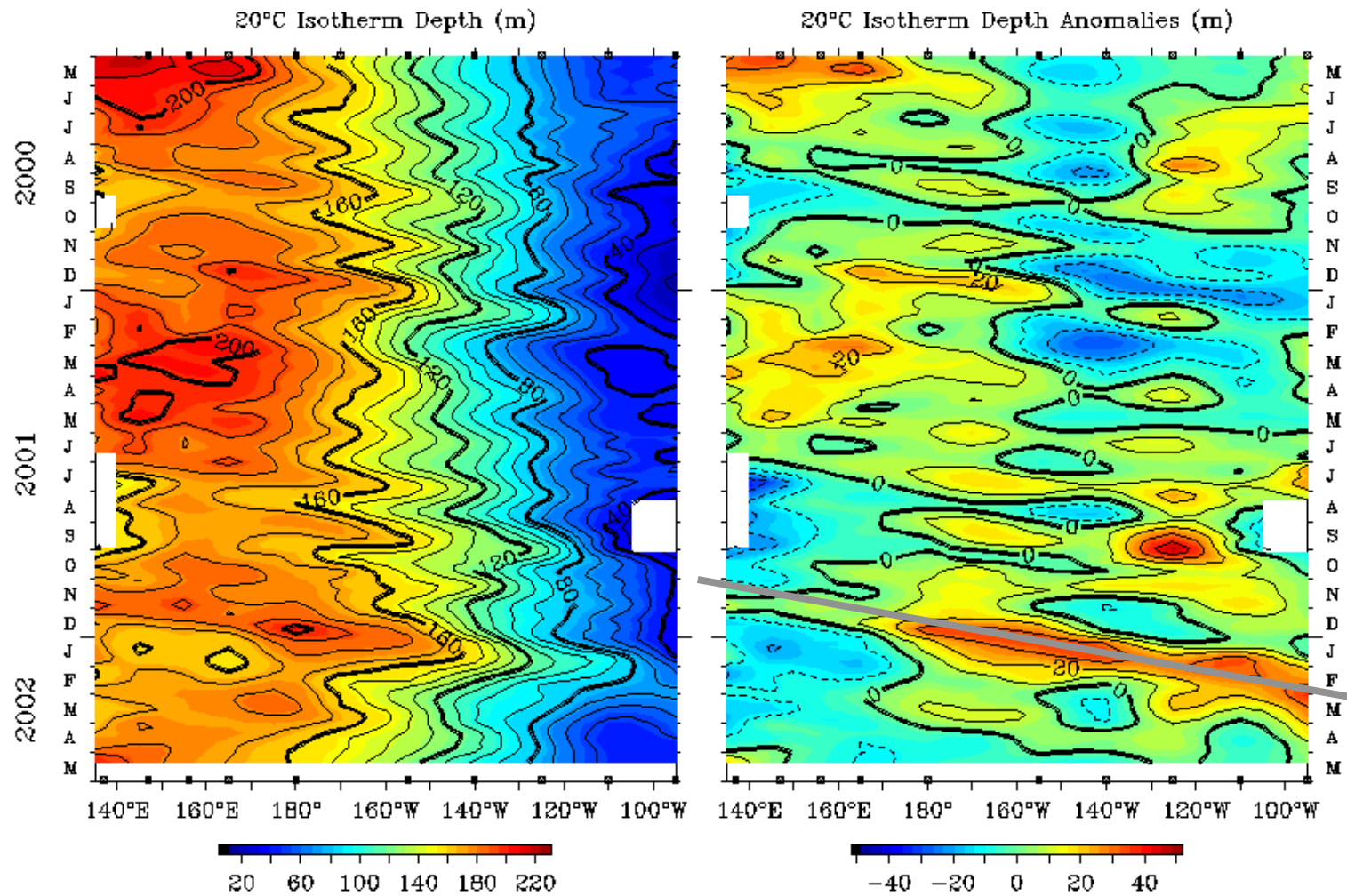


Equatorial Kelvin Waves

- Satellite altimetry from TOPEX/Poseidon
- Scenes are 10 days apart



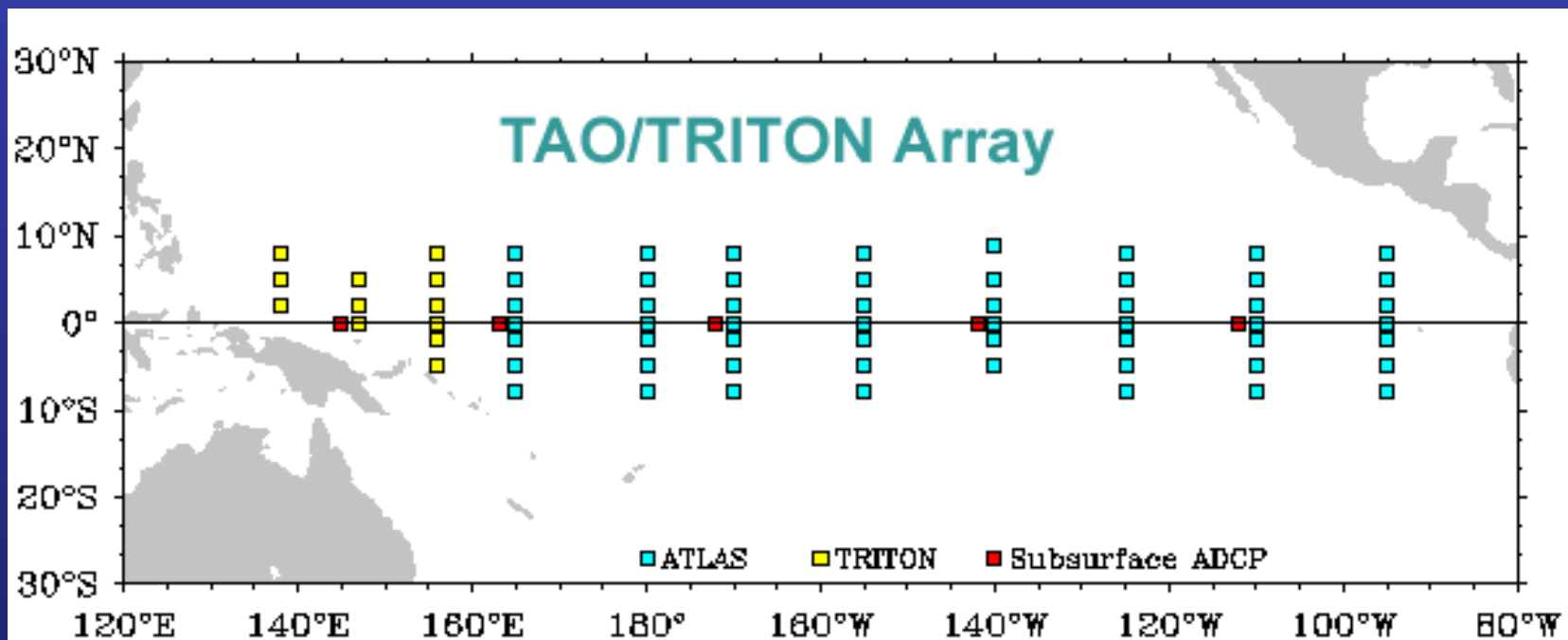
Five-Day 20°C Isotherm Depth 2°S to 2°N Average



Propagates 13,000 km in 4 months - 1.3 m/s

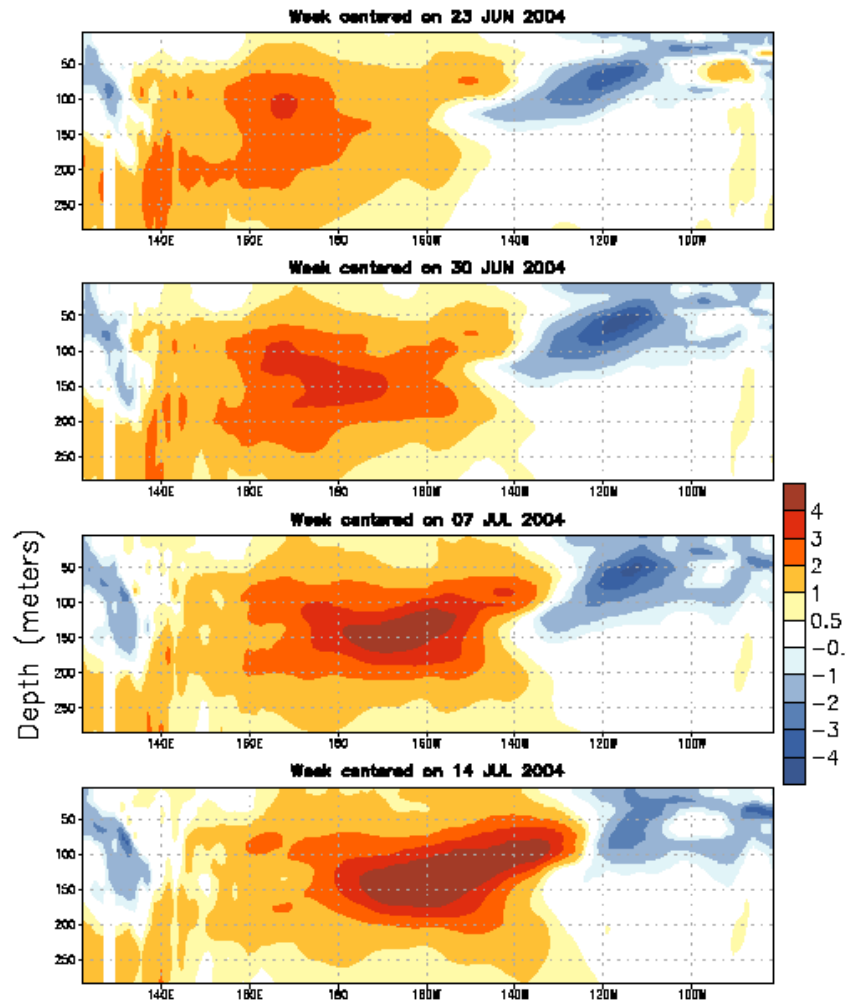
TOGA-TAO Array

- Equatorial array of buoys
- U.S., Japan & French partnership
- <http://www.pmel.noaa.gov/tao/>

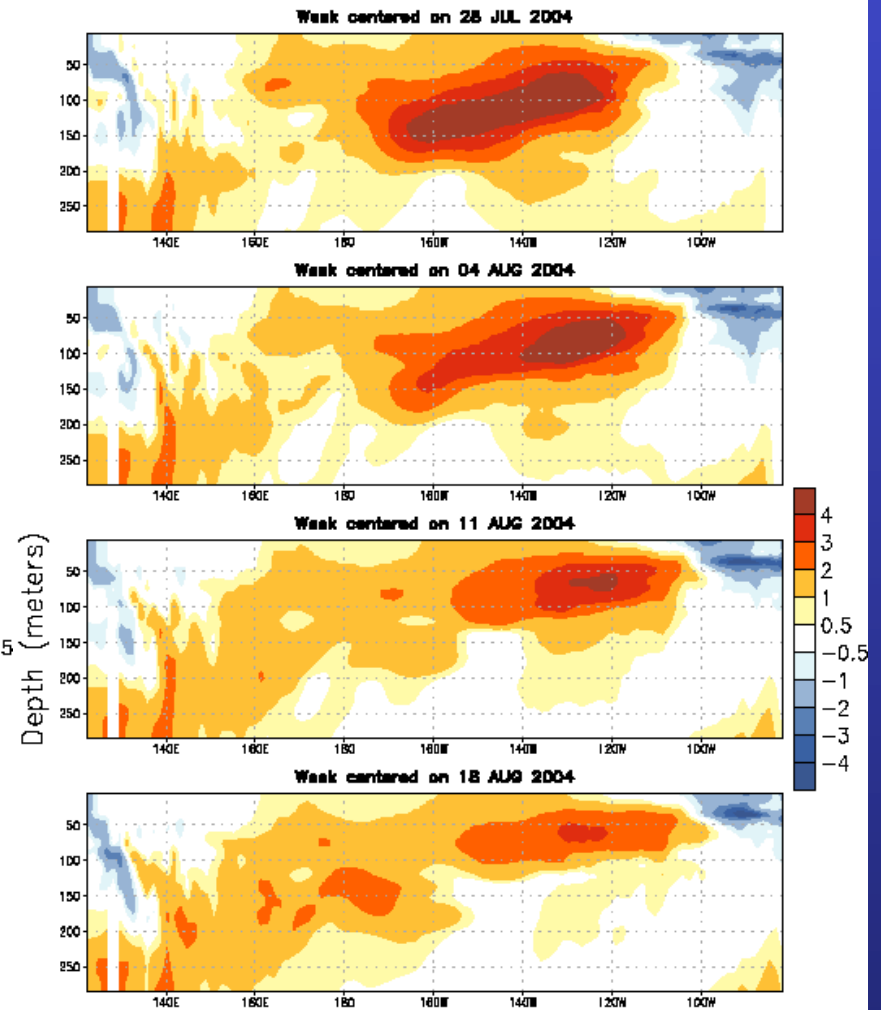


Equatorial Kelvin Waves

EQ. Subsurface Temperature Anomalies (deg C)

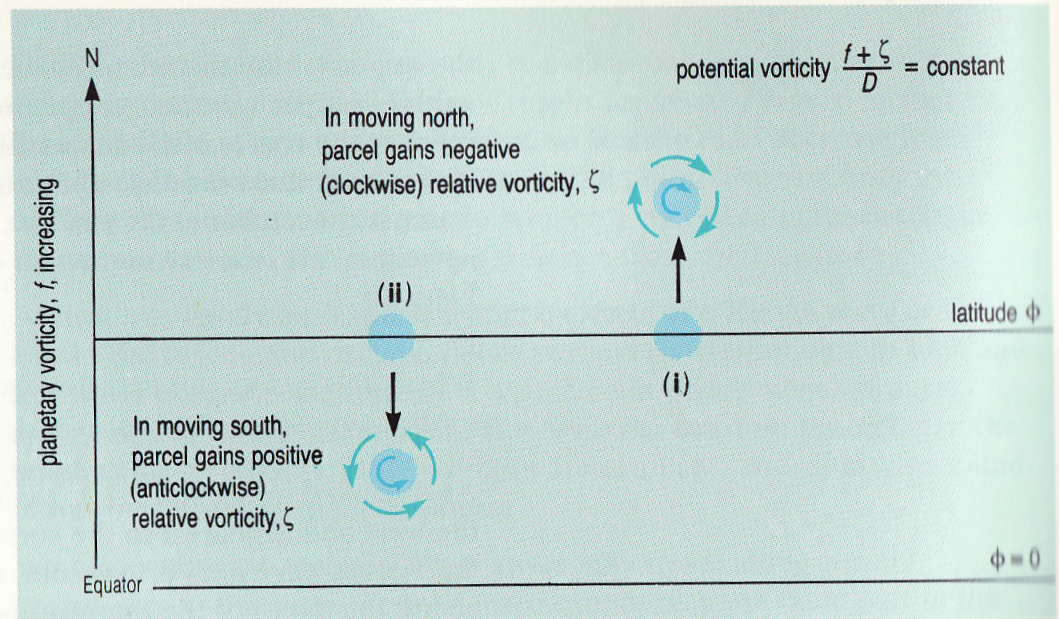


EQ. Subsurface Temperature Anomalies (deg C)

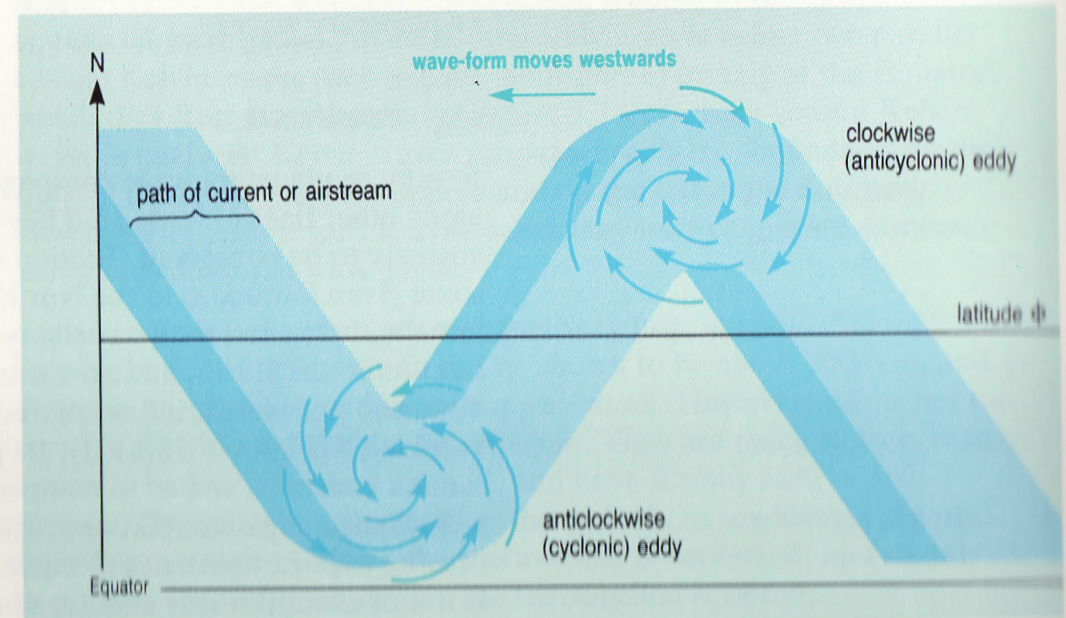


Rossby wave

- Restoring force = conservation of PV
- Motion in the horizontal plane



(a)



(b)

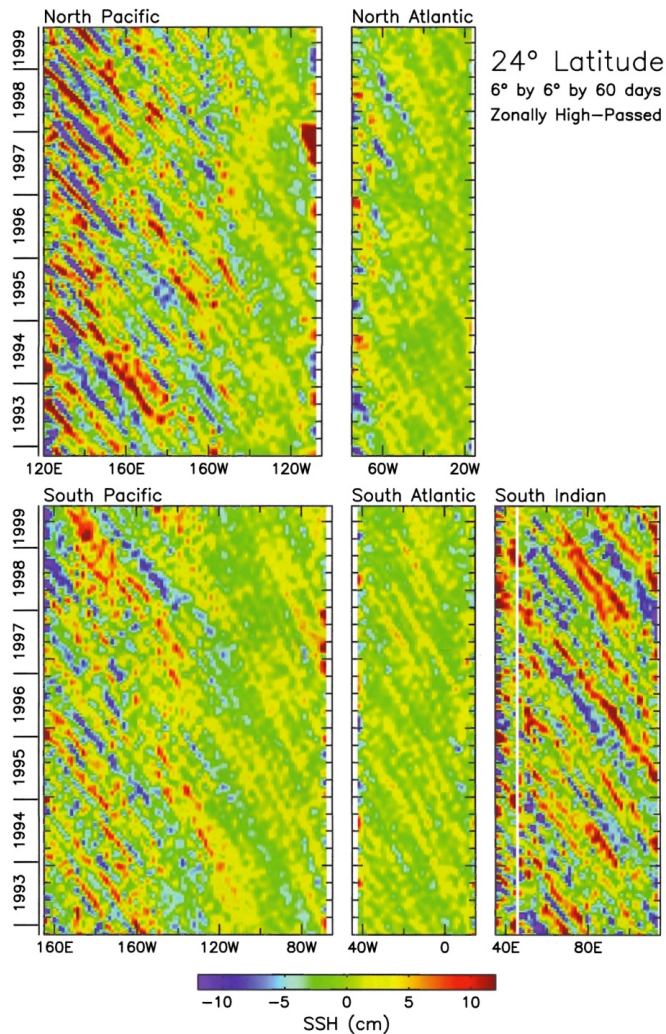


FIGURE 14.18

Surface-height anomalies at 24 degrees latitude in each ocean, from a satellite altimeter. This figure can also be found in the color insert. *Source: From Fu and Chelton (2001).*

TALLEY

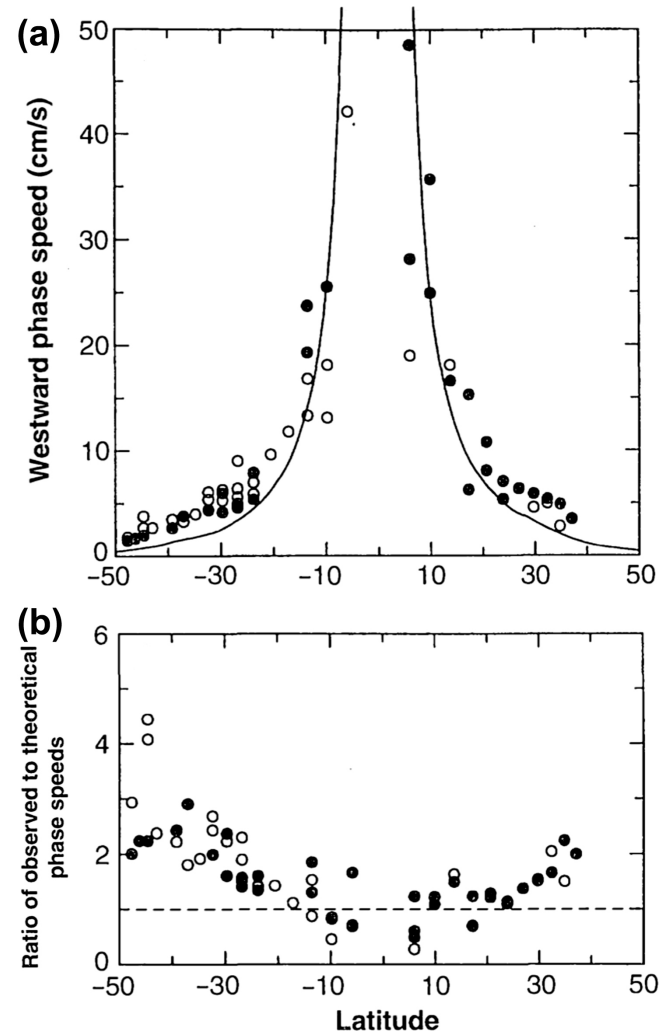
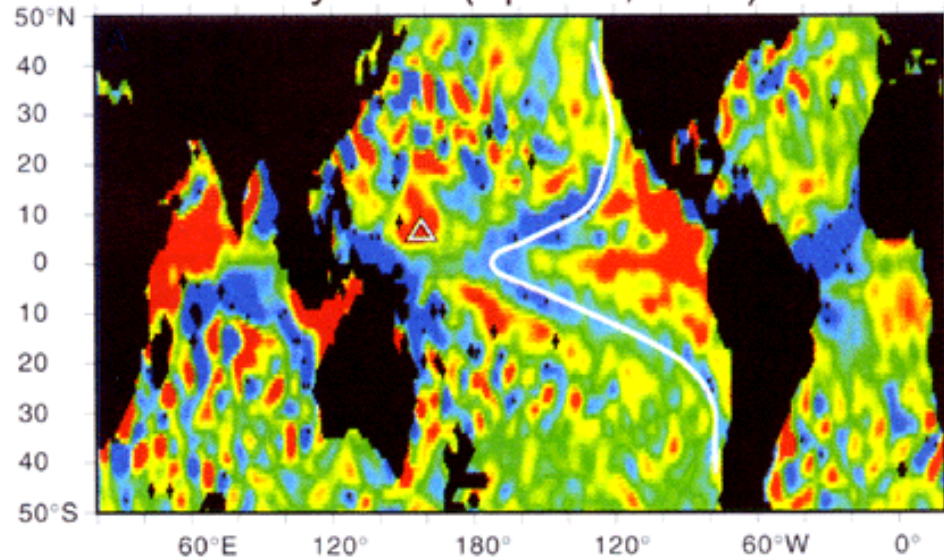


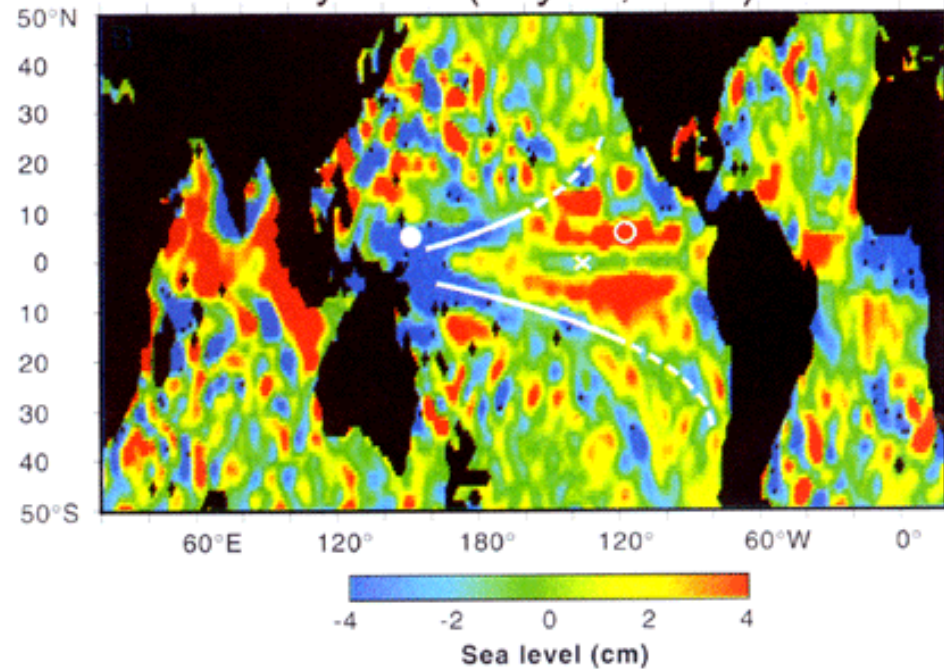
Figure 14.19 (a) Westward phase speeds (cm/sec) in the Pacific Ocean, calculated from the visually most-dominant SSH anomalies from satellite altimetry. The underlying curves are the fastest first-mode baroclinic Rossby waves speeds at each latitude. (b) The ratio of observed and theoretical phase speeds, showing that the observed phase speeds are generally faster than theorized. *Source: From Chelton and Schlax (1996).*

Rossby Waves

Cycle 21 (April 13, 1993)



Cycle 32 (July 31, 1993)



Waves critical to ENSO

- Panel 1: Internal equatorial Kelvin wave approaches (on 20C isotherm)
- Panel 2: Coastal Kelvin wave forms
- Panel 3: Reflected Rossby wave and coastal Kelvin

